

Dynasonde and VIPIR Ionosonde

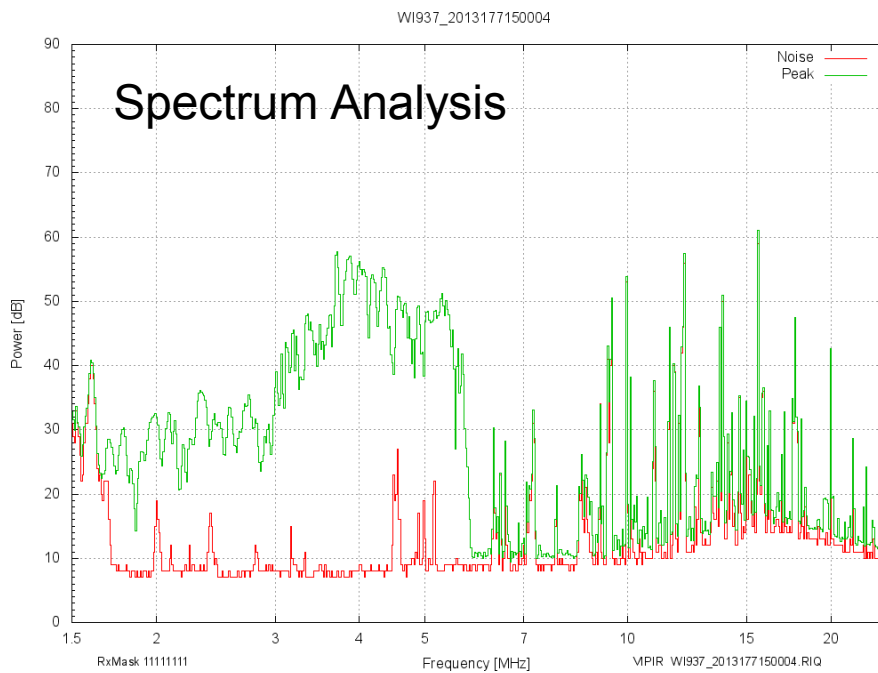
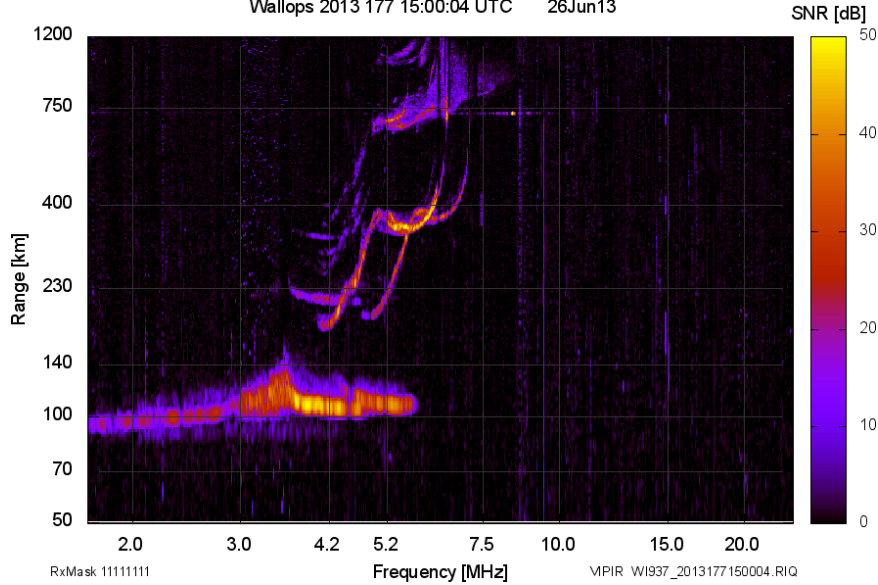
Field Site Requirements Basic and Research Capabilities

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with
NOAA
National Geophysical Data Center
Terence.Bullett@colorado.edu

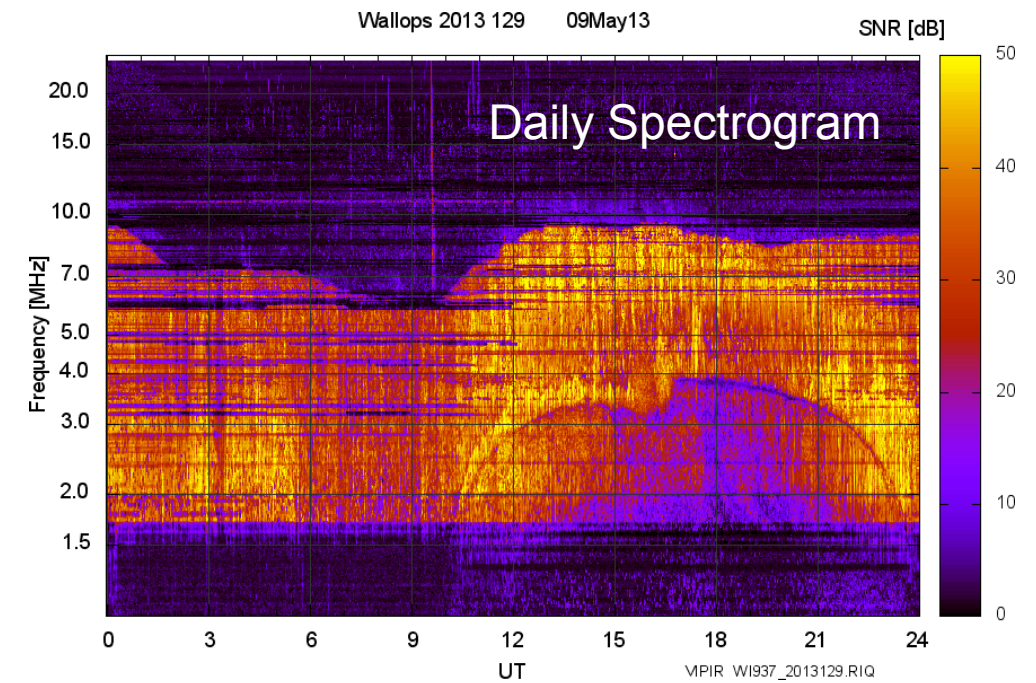
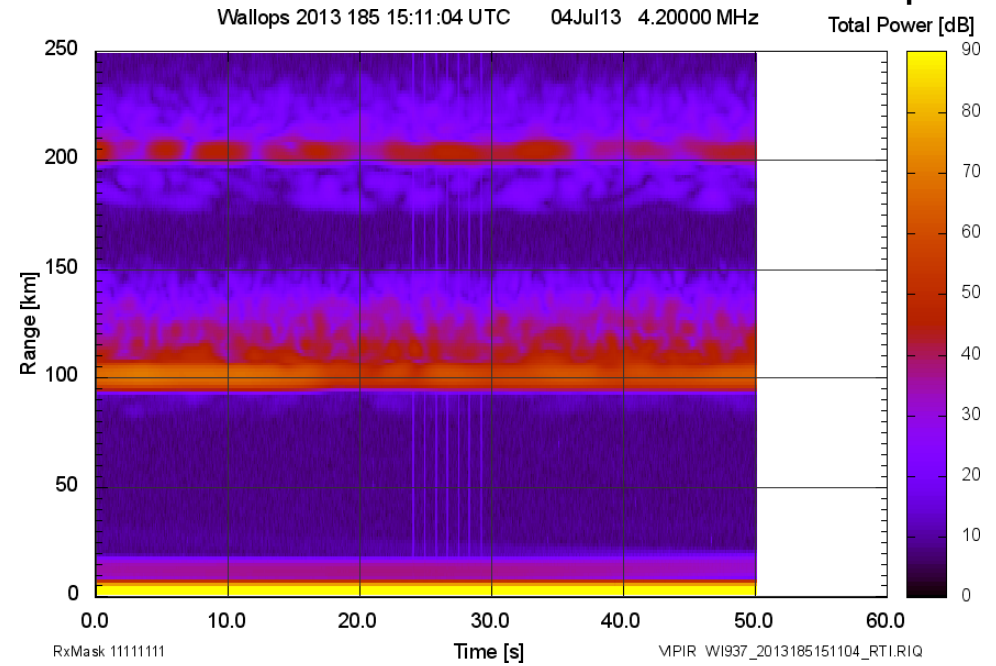
VIPIR is the Hardware

SNR Ionograms

Wallops 2013 177 15:00:04 UTC 26Jun13

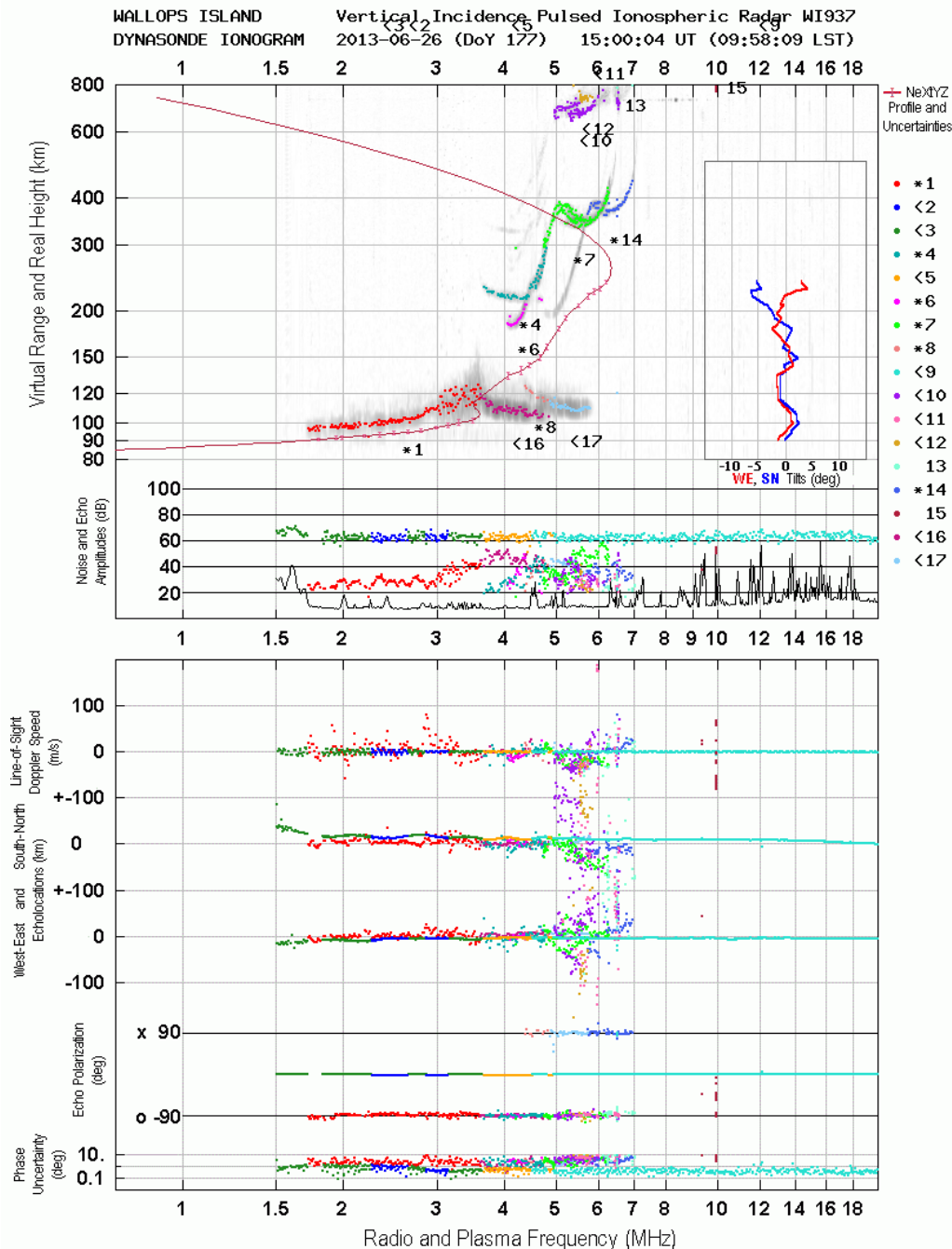


Fixed Frequency



Dynasonde is the Software

processed automatically



- Echo Detection
- Trace Classification
- Ionogram Scaling
- 3D echo-location
- Electron Density Profile
- Tilt Profile

Vertical Incidence Pulsed Ionospheric Radar

Very high interference immunity: $IP3 > 45 \text{ dBm}$

High Dynamic Range: $115(I) + 30(V) \text{ dB}$

Direct RF sampling 14 bits at 80 MHz

Fully digital conversion, receiver and exciter

Waveform Agility: $2 \mu\text{s}$ to 2 ms pulse/chip width

USB-2 Data and Command/Control Interfaces

8 coherent receive channels; Frequency: $0.3 - 26 \text{ MHz}$

4 kW class AB pulse amplifier: 3^{rd} harmonic $< -30 \text{ dBc}$

Precise GPS timing possible for bi-static operation

Radar software Open Source C code; runs under Linux

Upgrades for the VIPIR Mark II

- FPGA based digital receiver
- 16 bit, 120 MHz ADC
- USB3 data transfer
- Improved analog front end
- Improved receive antenna pre-amplifiers
- Contemporary computers and data storage
- Options:
 - High power low pass transmit harmonic filter
 - Rubidium oscillator for oblique phase measurements

VIPIR Mark II



Jang Bogo Antarctic Research Station
Korean Polar Research Institute

Wallops Island VIPIR hardware



Power Conditioner

4kW RF Amplifier

KVM

Exciter

Reference

Receiver

Front End

Balun

Control Computer

Analysis Computer

UPS

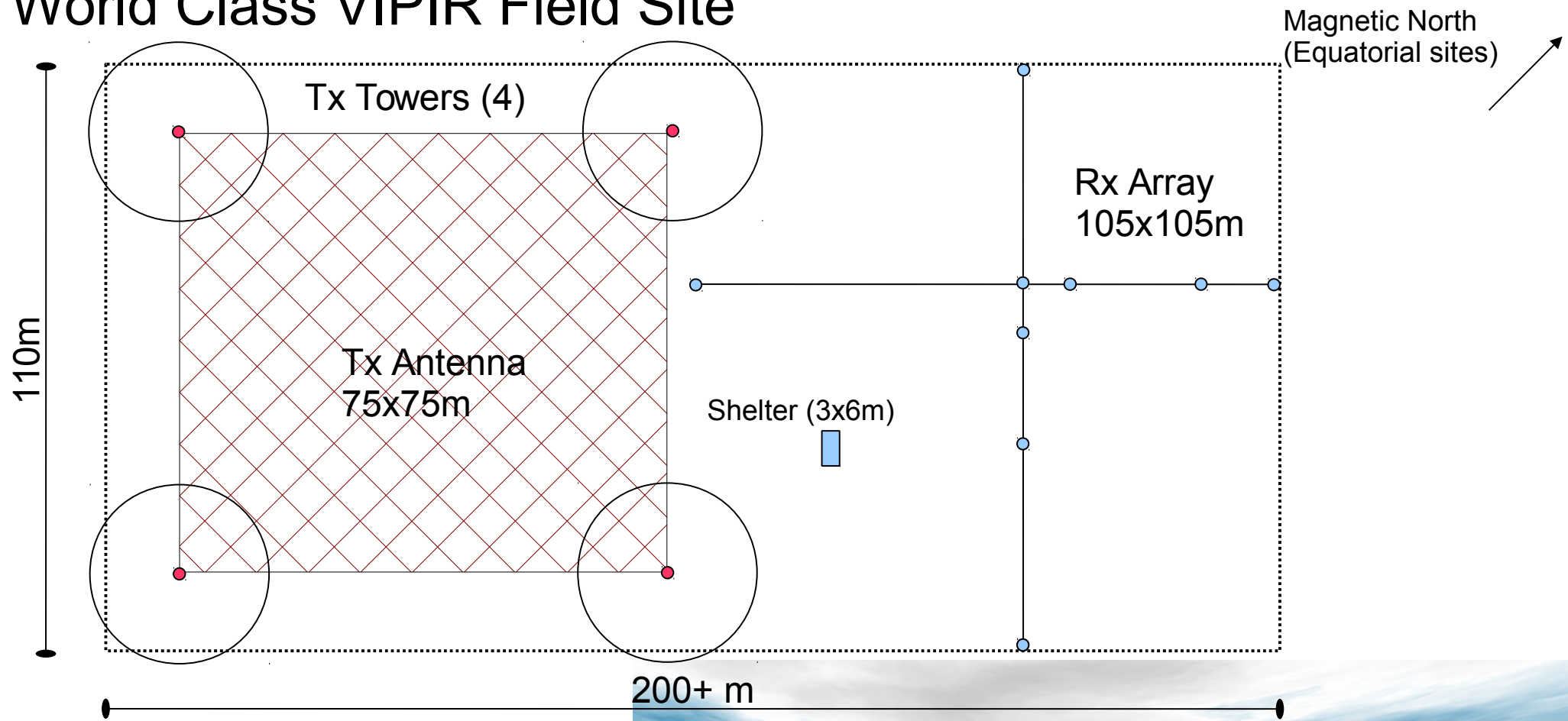
Standard 19" rack, 46" tall

Installation “classes”

- Research
 - Major research facility for plasma physics, propagation research and ionosphere discovery. Performance is primary, cost and footprint are secondary concerns.
- Average
 - Ionosphere monitoring and geophysical research. A compromise between of cost and capability.
- Small
 - Ionospheric monitoring. Cost and footprint are primary concerns and limitations.
- Transportable
 - Ionospheric measurements from various locations. Mobility is the primary concern.

There is approximately a factor of 10 in performance between these classes

World Class VIPIR Field Site



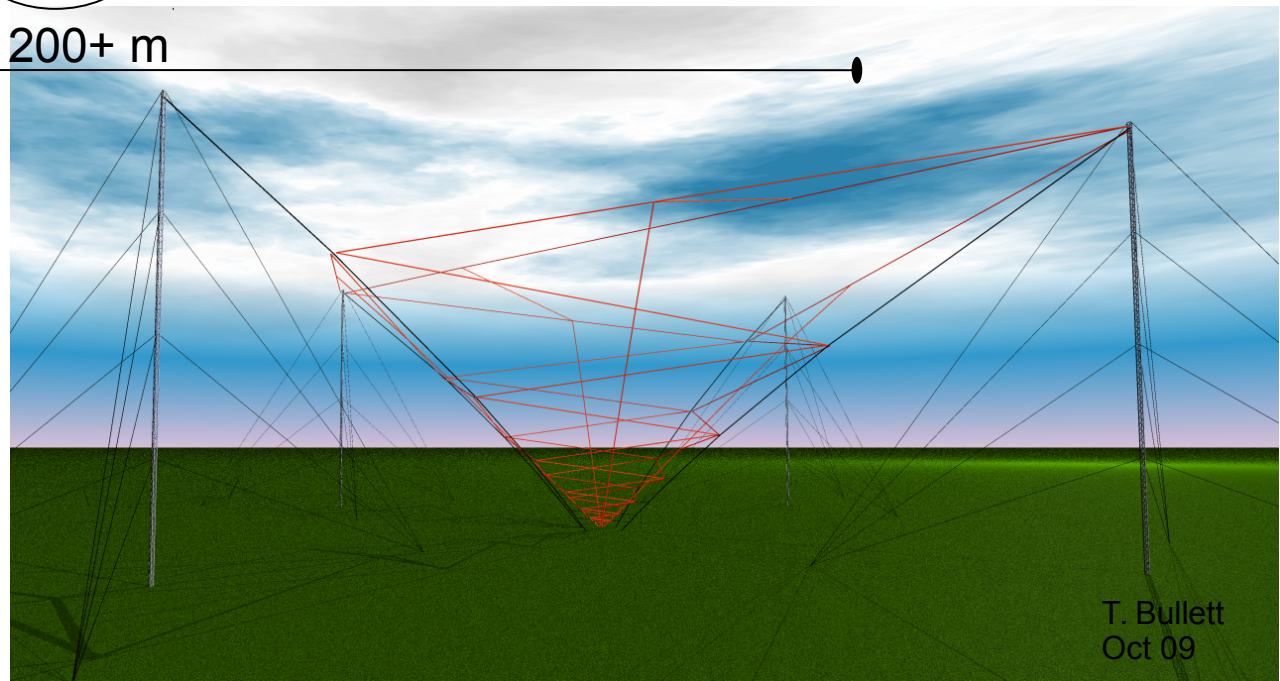
8 channel HF radar

8 or 10 element Rx array

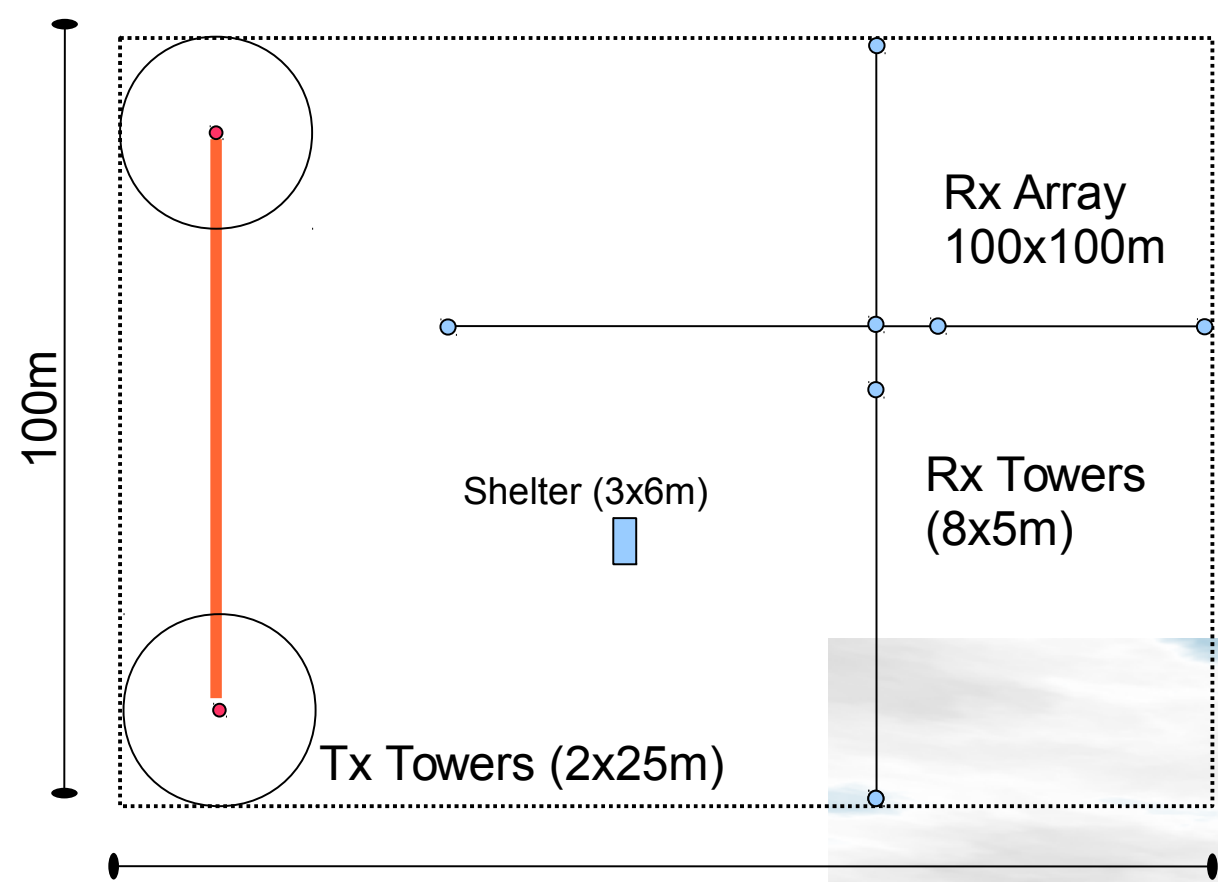
Log-periodic Tx antenna

35m steel towers (4)

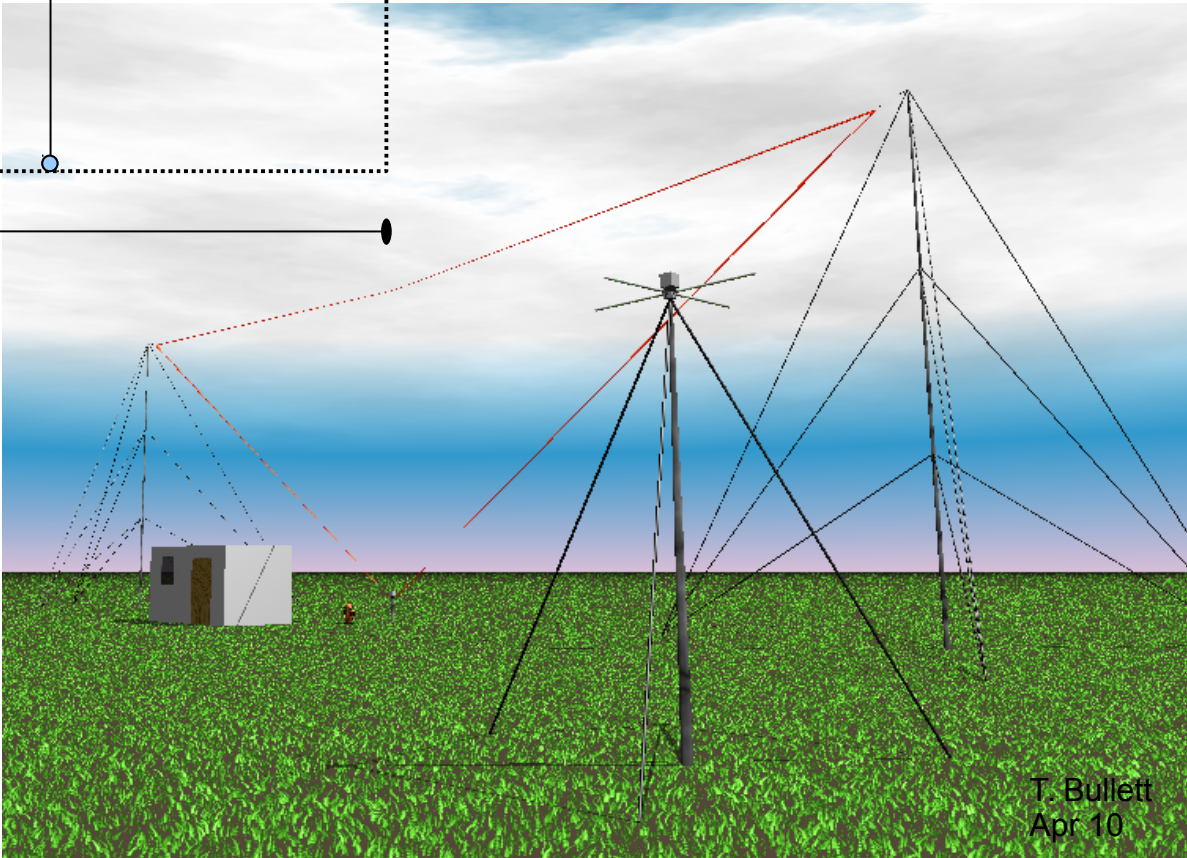
5 to 10 acre footprint



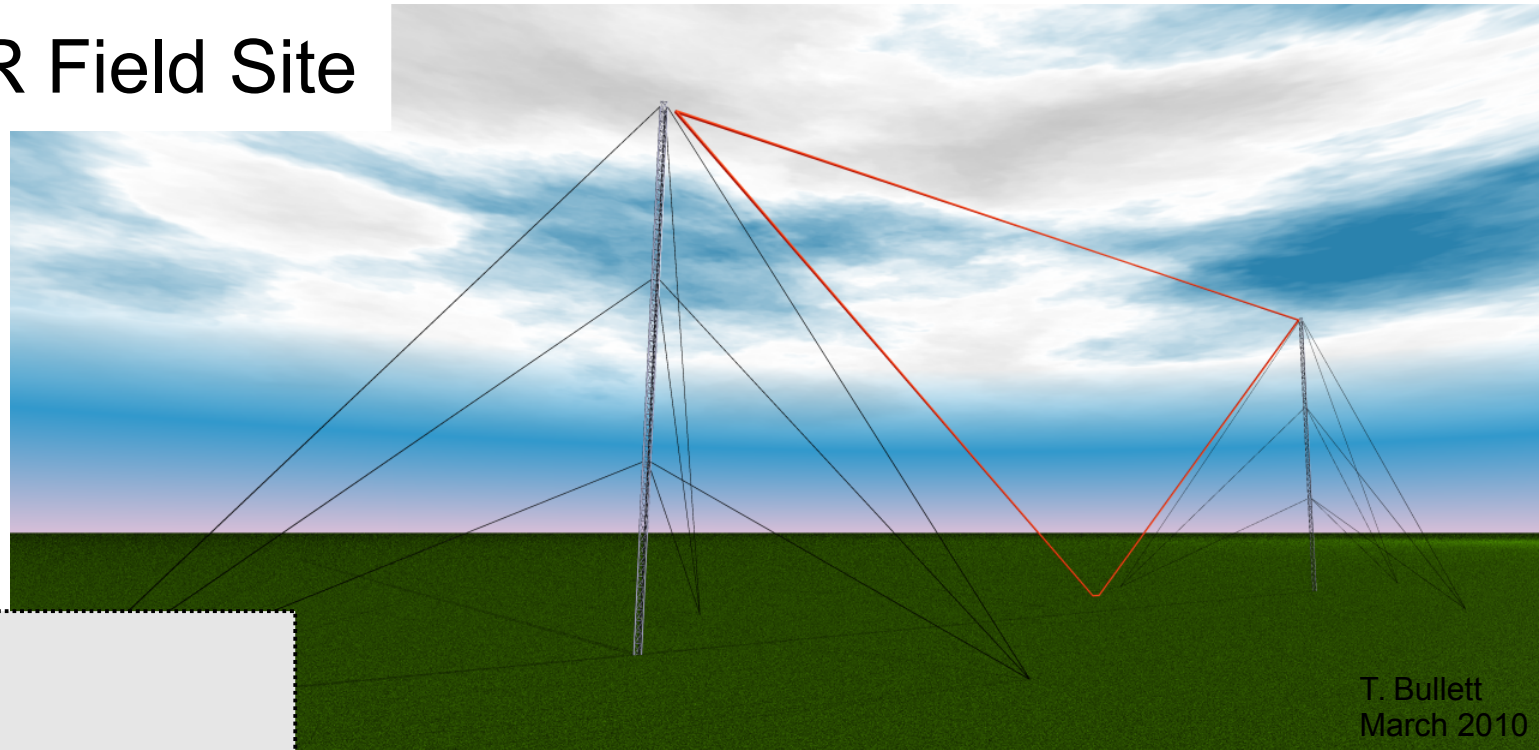
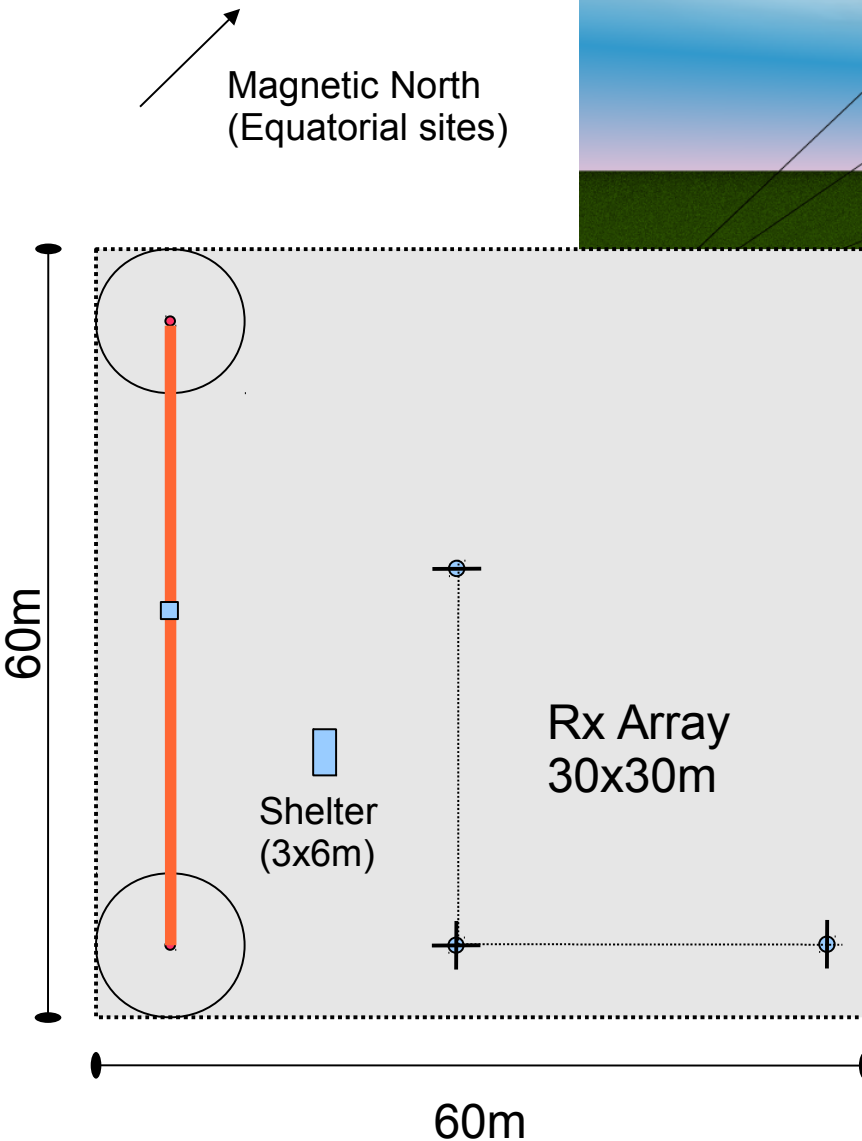
Intermediate VIPIR Field Site



- 8 channel HF radar**
- 8 element Rx array**
- Inverted Delta Tx antenna**
- 25m Tx Towers (2)**
- 5m Rx towers (8)**
- 3.5 acre footprint**



Reduced VIPIR Field Site



4 channel HF radar

4 element Rx array

Inverted Delta Tx antenna

15m Tx poles (2)

5m Rx poles (4)

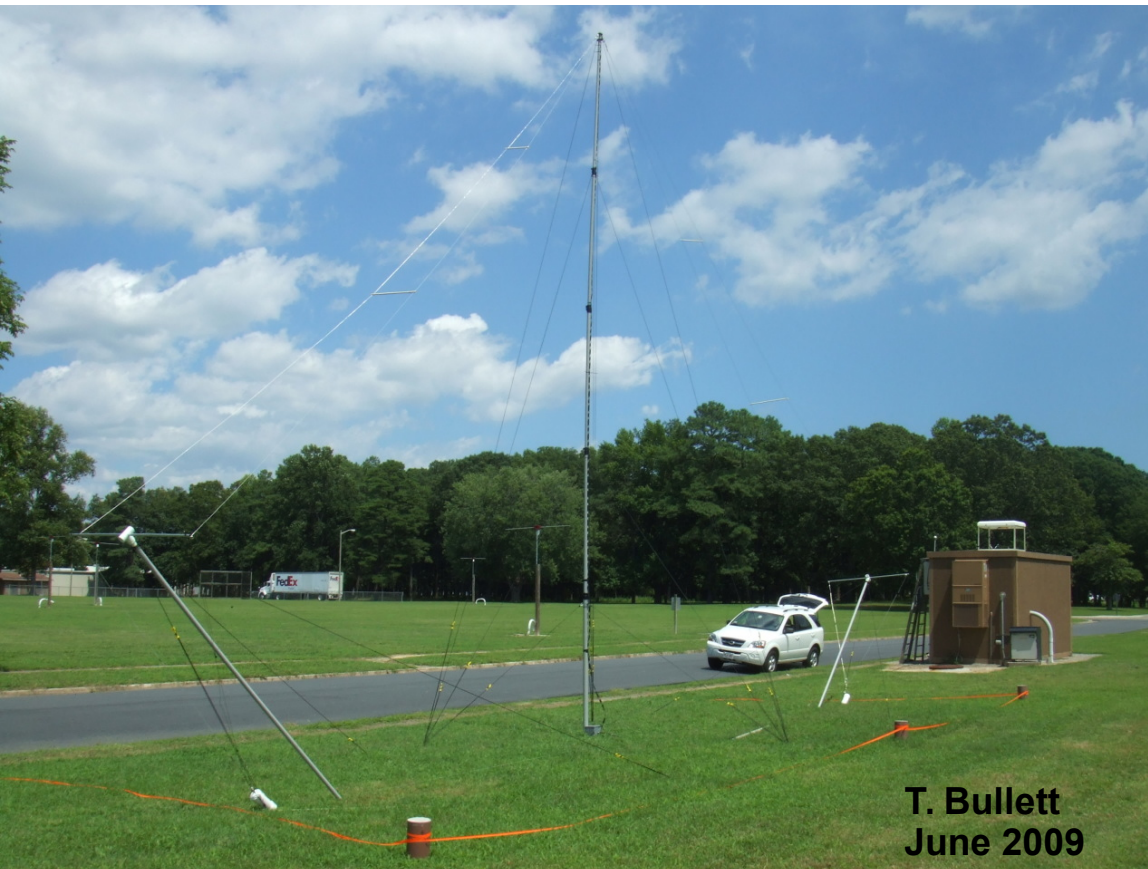
< 1 acre footprint

Virtual walk-through:

http://www.ionosonde.com/images/Ionosonde_Tour.avi

Transportable VIPIR

Photographs of VIPIR prototype antennas designed for use on a 40 ft fishing boat.



Portability is key

Antennas engineered for temporary install & transport

4 or 8 channel HF radar

2-8 element Rx array

Mini-Delta Tx antenna or custom

Facilities

Climate Controlled Shelter

3mx3m minimum

Larger space enhances research

25,000 BTU Air Conditioner (Large window model)

220v x 15A or 110Vx30A power

Internet connection

128kbps minimum ; 1Mbps recommended

Absence of:

High Voltage Power Lines

Industrial electrical noise

Other HF users (receive and transmit)

MF transmitters (AM Broadcast or NDB)

Local Construction

Shelter

Receive anchors

Receive poles

Transmit anchors

Transmit towers

RF Cable conduit

Cable Vault

Your existing Tx
antenna!

Local Parts (Optional)

19" equipment rack

Rack Slides

Power Conditioner

UPS

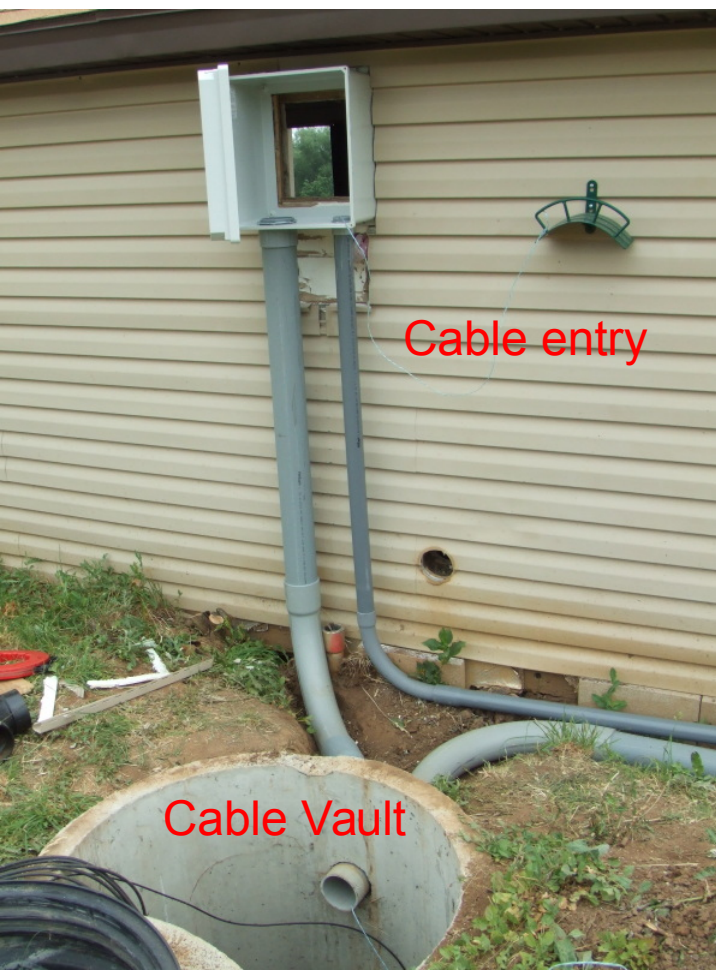
Keyboard

Mouse

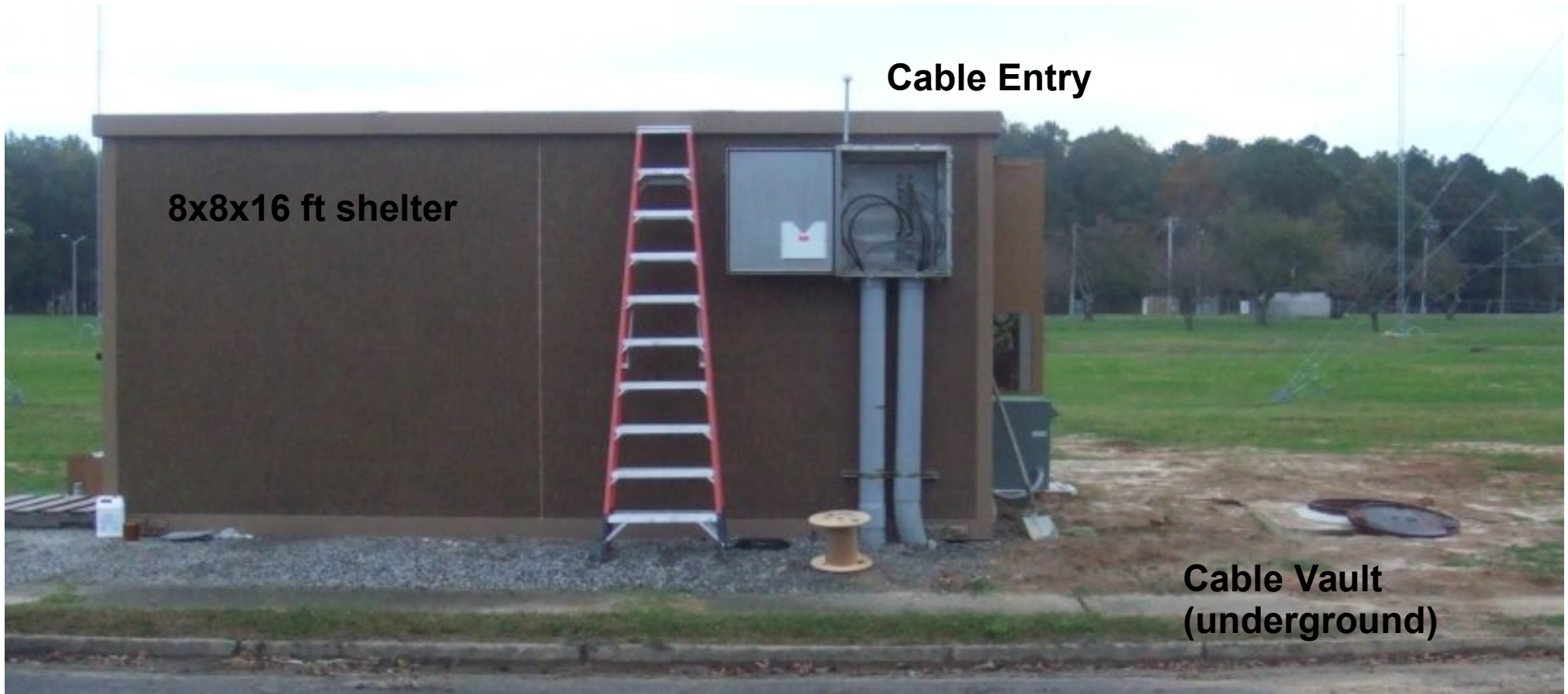
Monitor

KVM switch

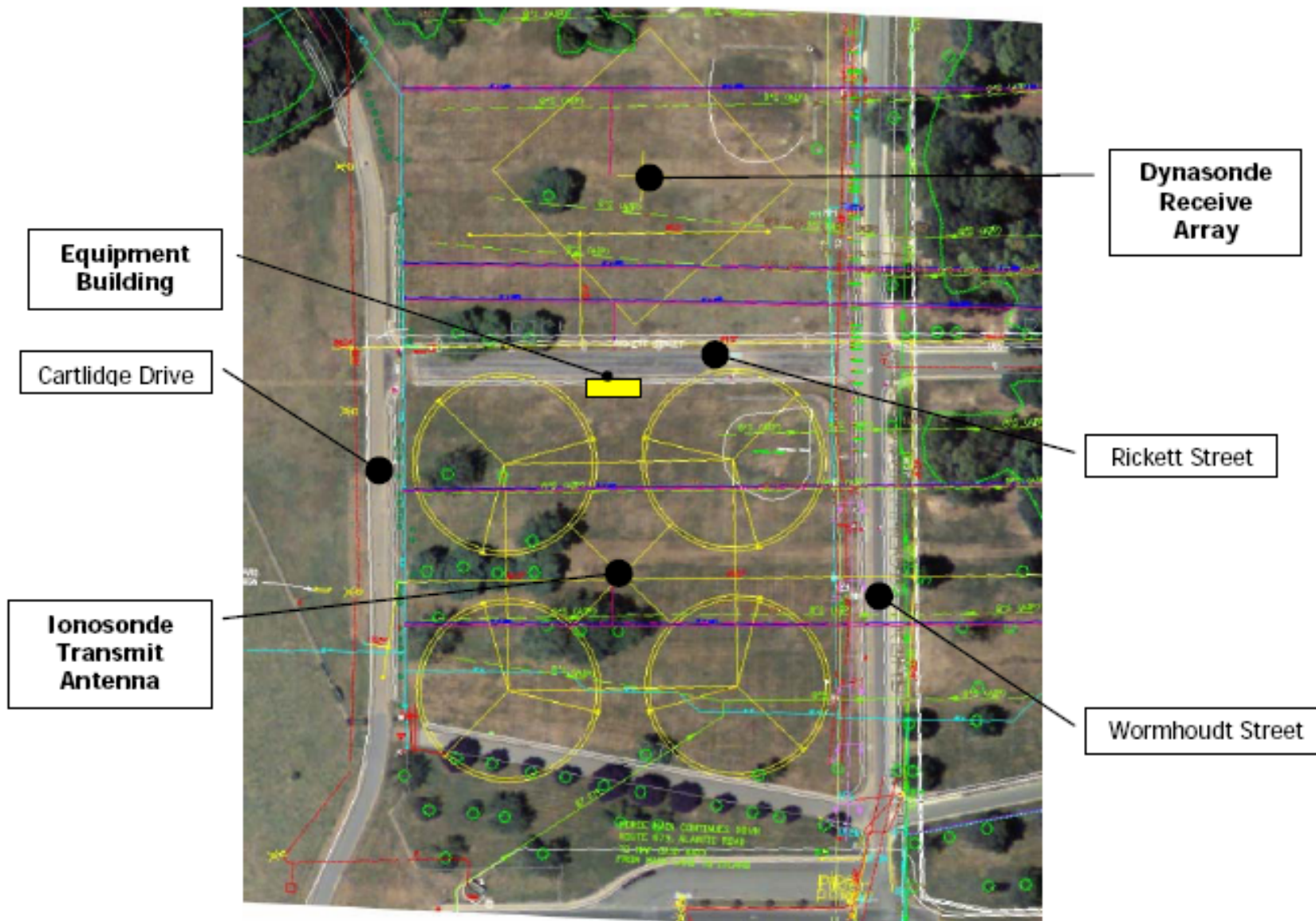
Boulder Site Photos



Wallops Shelter



Wallops VIPIR Field Site



Long Term “World Class Observatory” being provided by NASA for continued research and routine observations

Wallops Island: September 2006

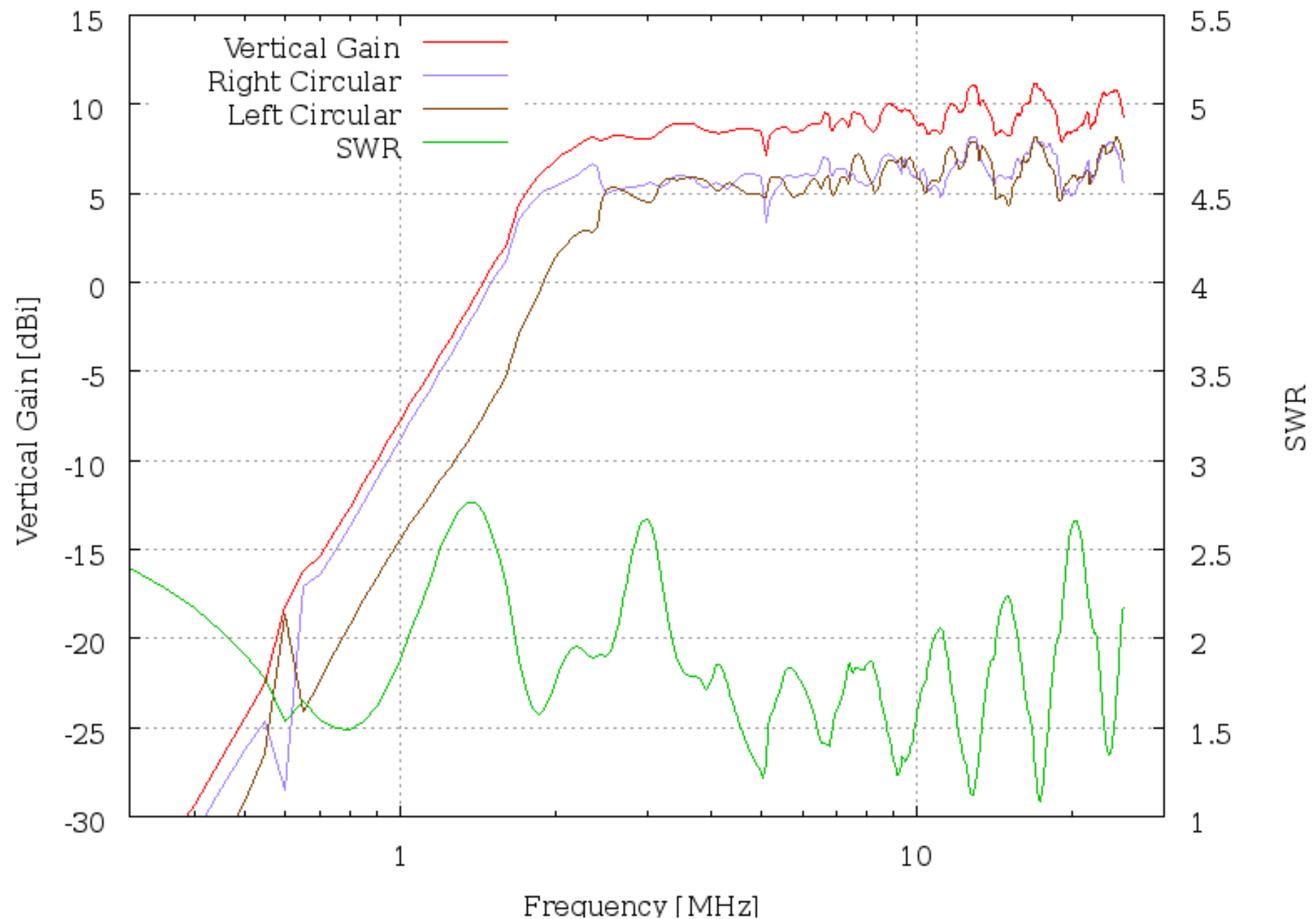


Wallops Island Tx antenna

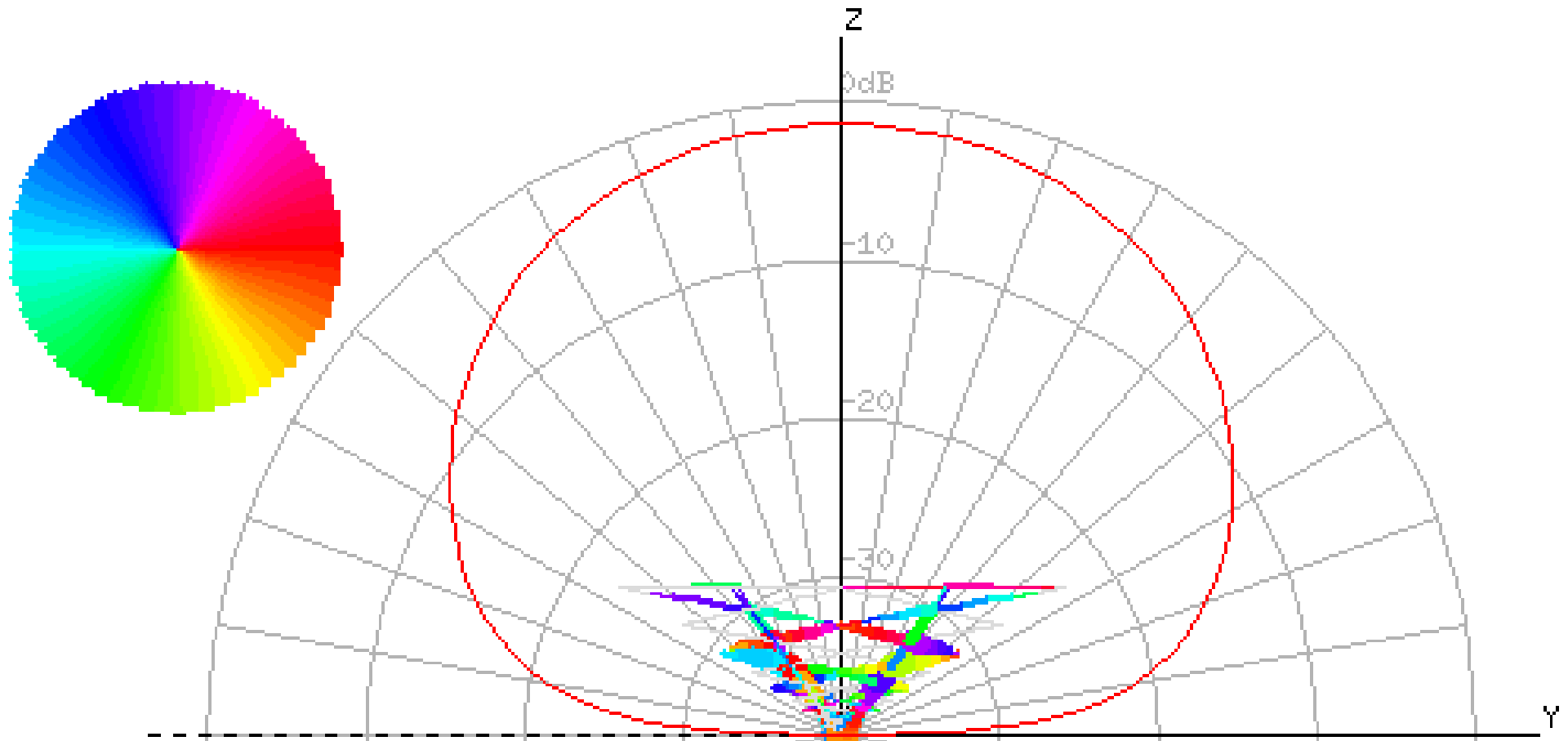


Transmit Antenna Performance

Wallops ZZLPA FOM: 29.054 7.681 5.990



Typical ZZLPA Pattern



Inverted Delta Transmit Antenna

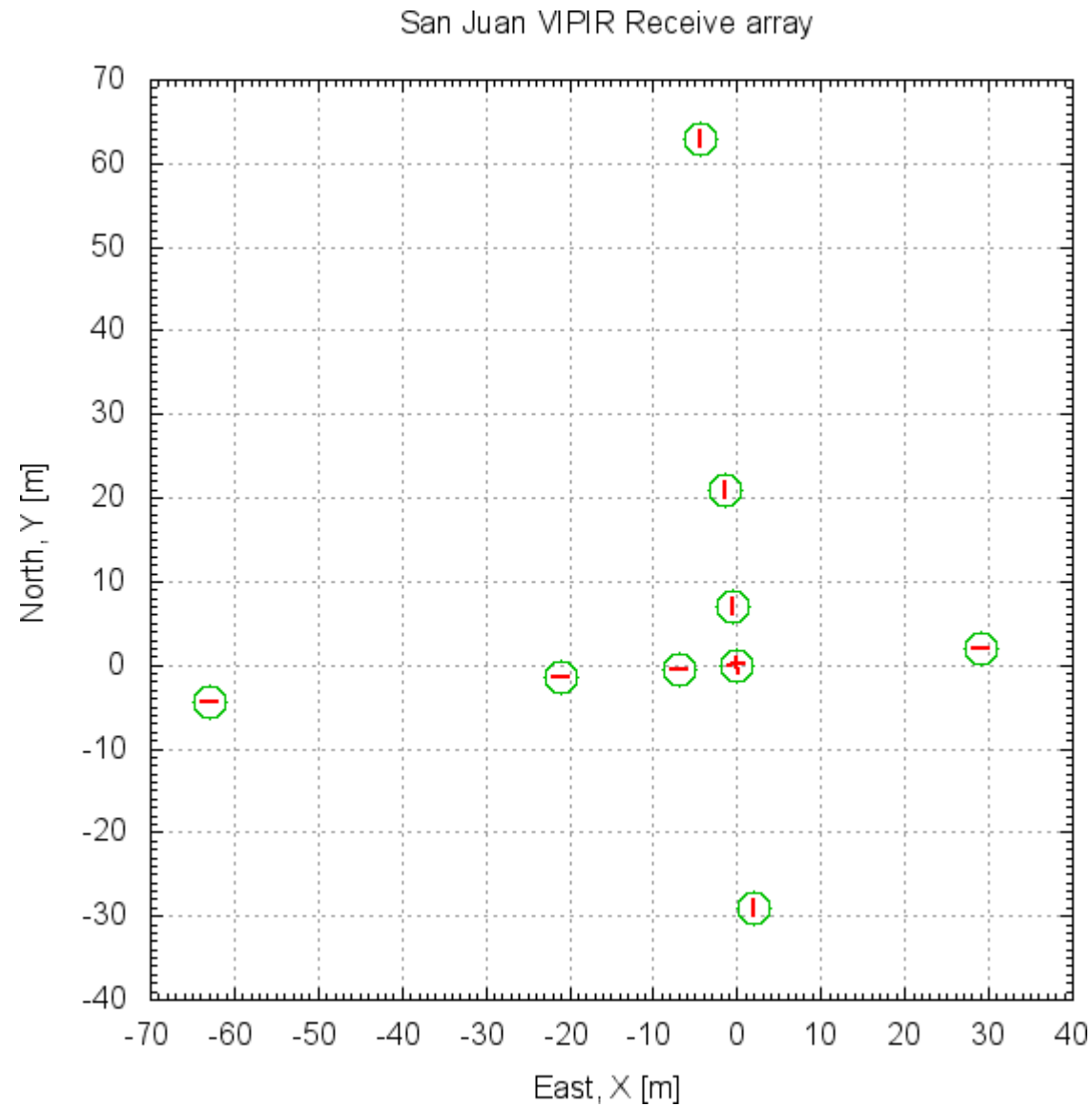


San Juan Observatory
“Small” : 15m tall x 45m long

Receive Antennas



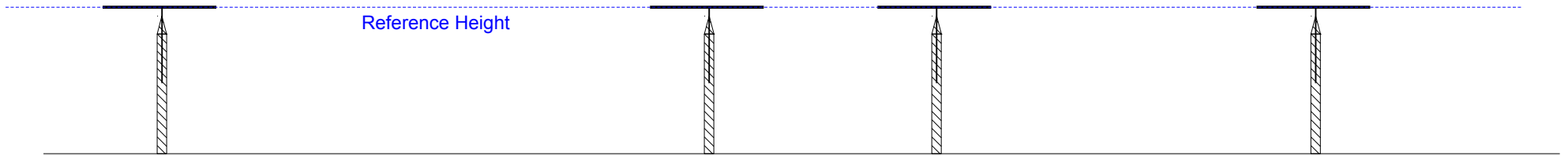
2m crossed dipoles



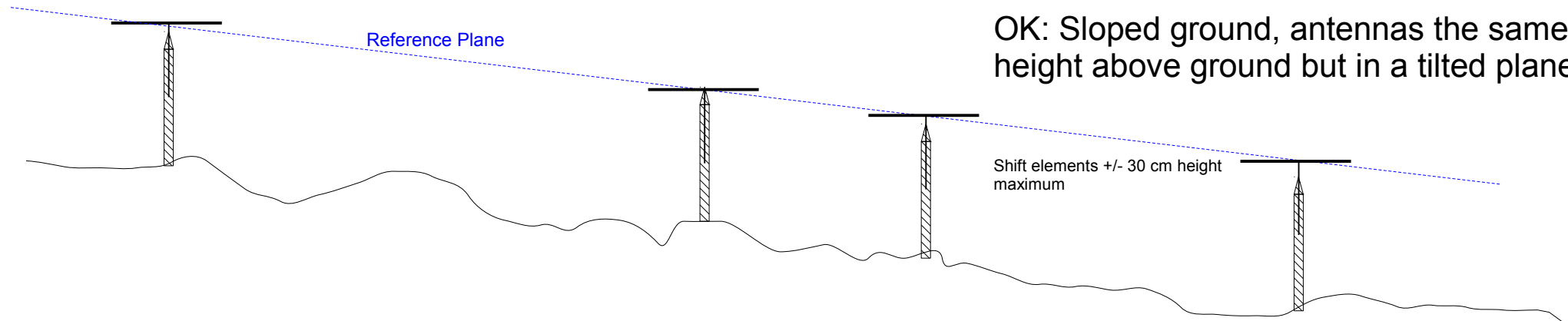
Antennas at locations -9 -3 -1 0 +4
Gives separations of 1 2 3 4 5 6 7 8 9 13
Units of 7m

Uneven Ground

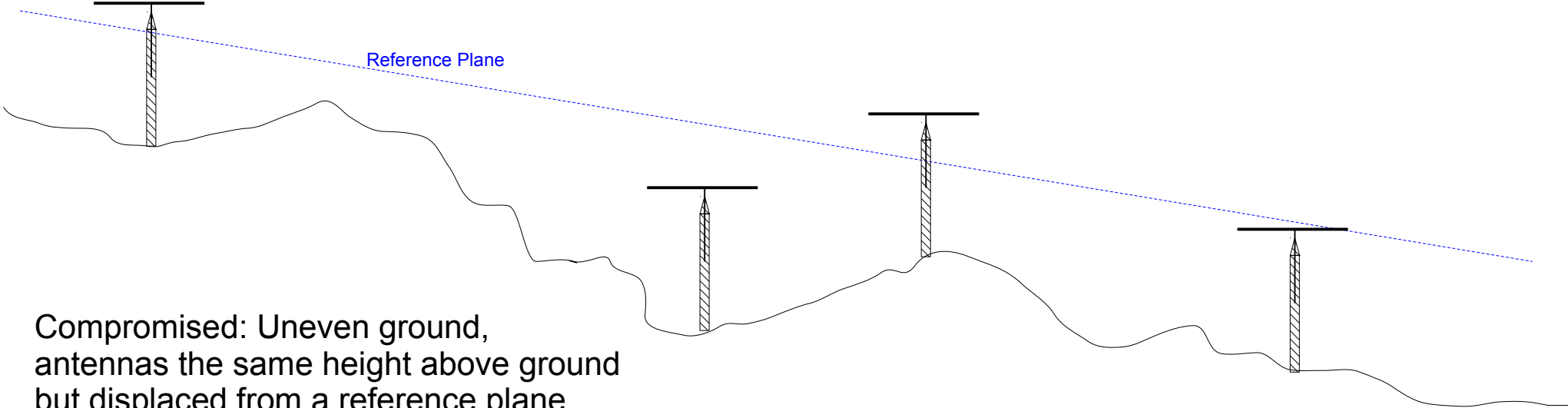
Ideal: Uniform, Horizontal ground, antennas the same height above ground



OK: Sloped ground, antennas the same height above ground but in a tilted plane



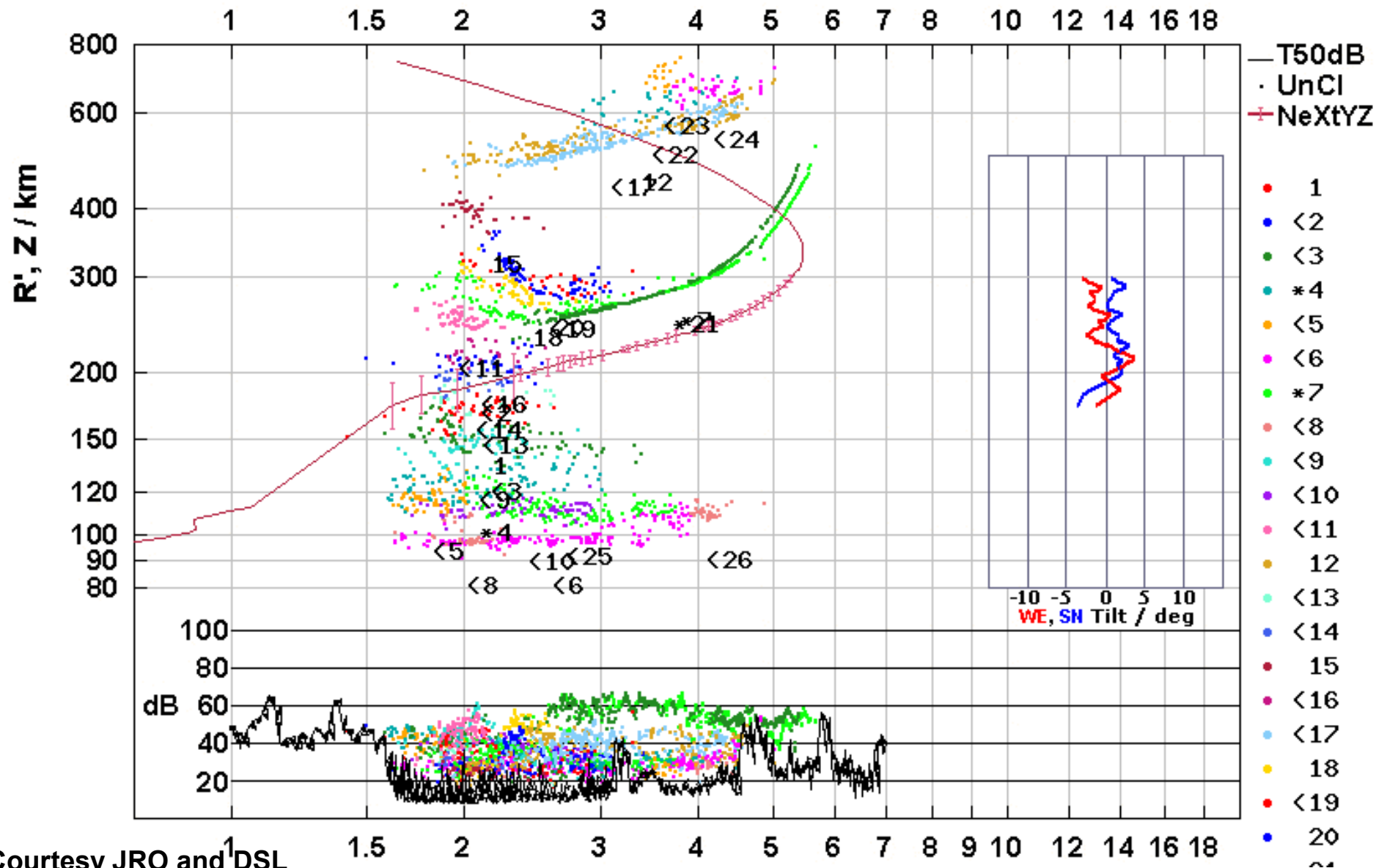
Compromised: Uneven ground, antennas the same height above ground but displaced from a reference plane



Ionogram Analysis: Dynasonde21

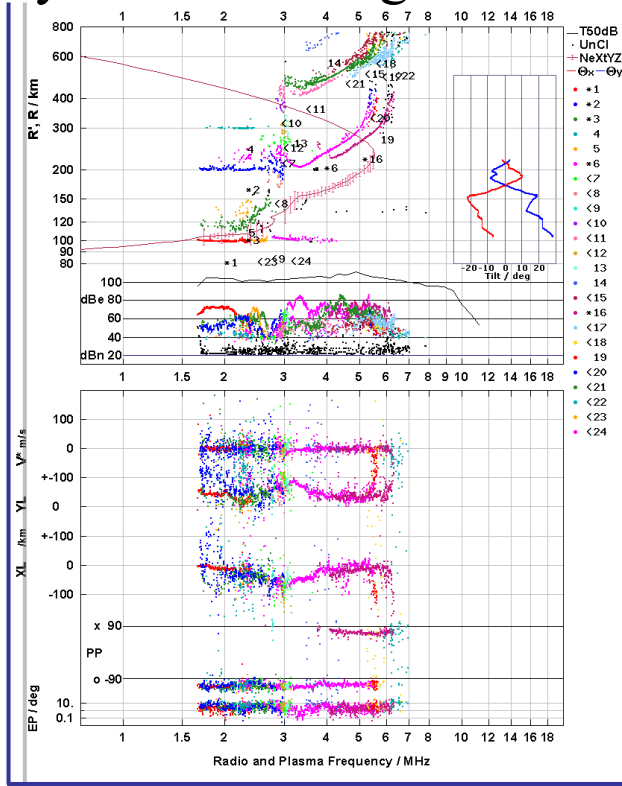
Jicamarca, Peru : VIPIR ionogram : Dynasonde21 Analysis

JICAMARCA OBS. DYNASONDE 08-08-06 2237UT 75°W
8pBP4It DoY= 219.942 File= RIQ.11111111

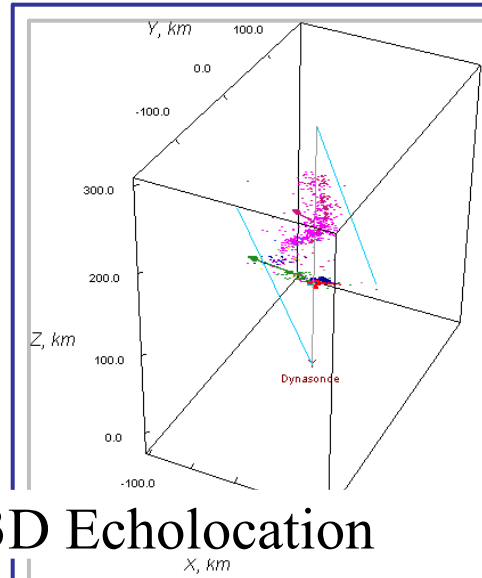


Dynasonde21

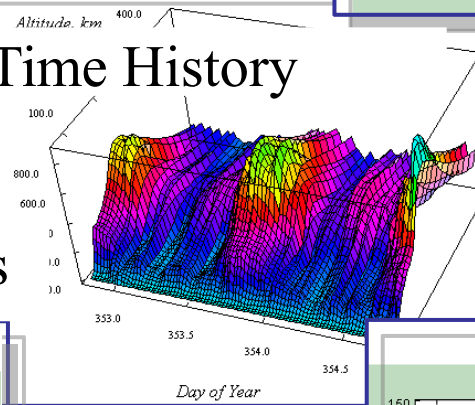
Dynasonde Ionogram Analysis



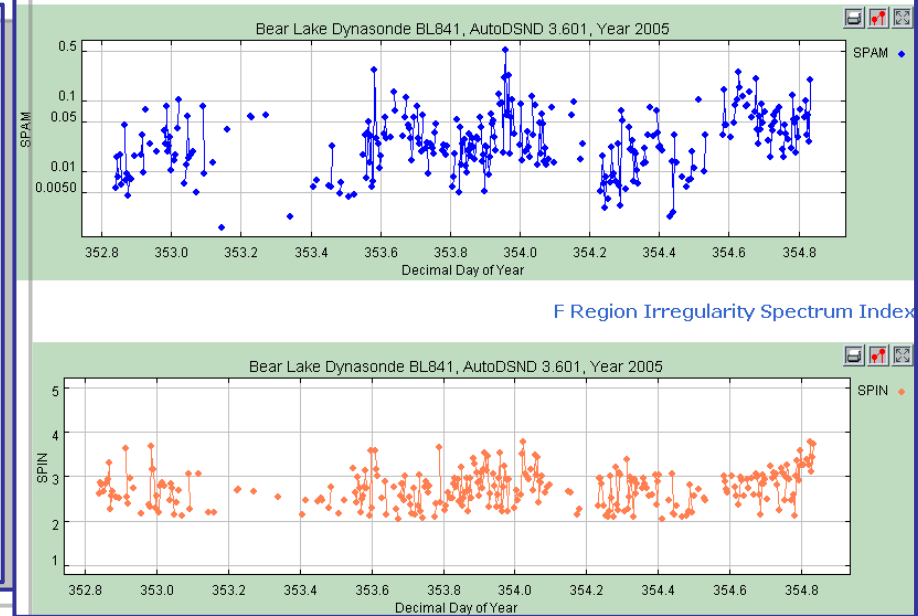
3D Echolocation



EDP Time History

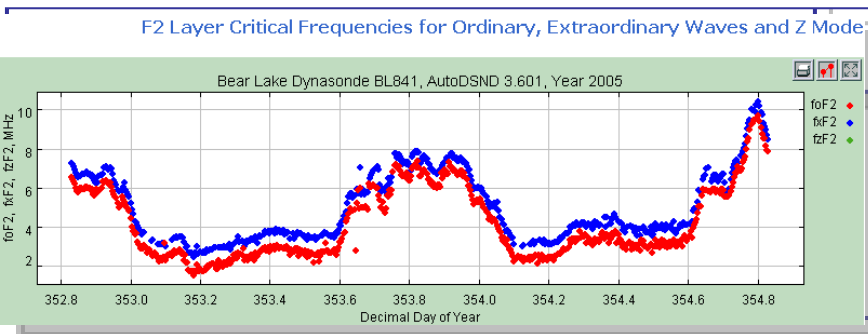


Structure Functions

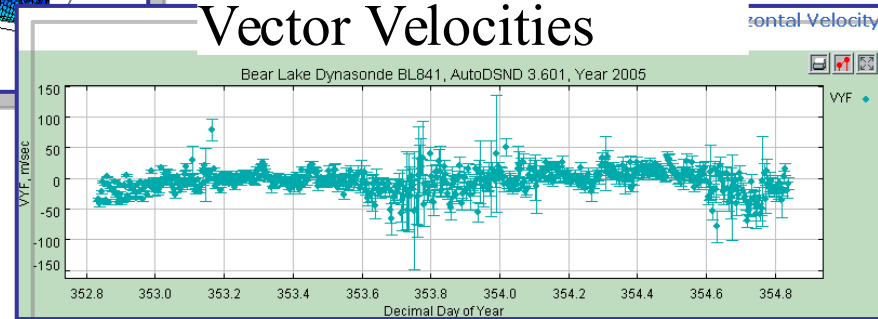


Dynasonde21 ©

Classical URSI scaled characteristics



Vector Velocities



Dynasonde is the name for methodology of ionospheric radio sounding based on the physical notion of radio echo and on a comprehensive use of phase information in it.

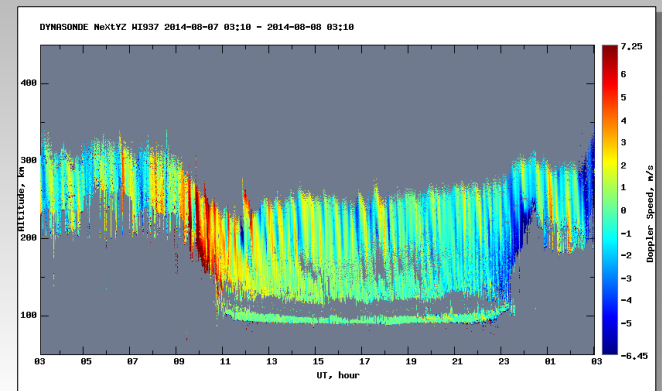
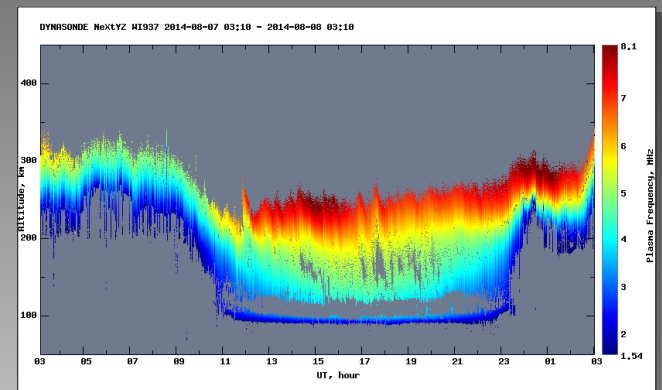
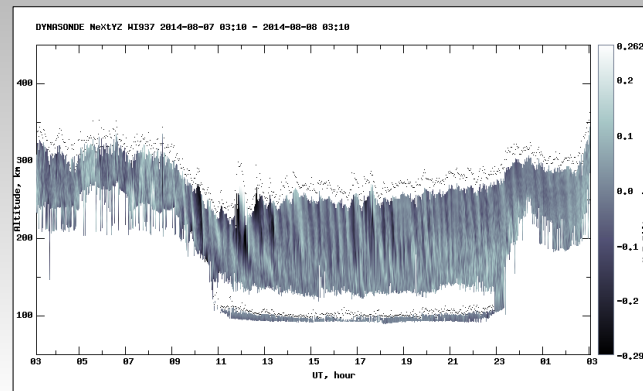
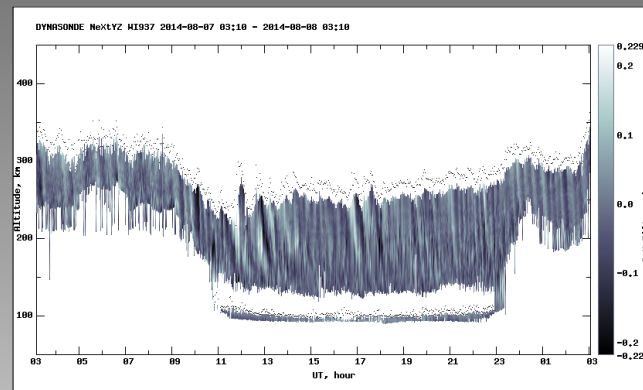
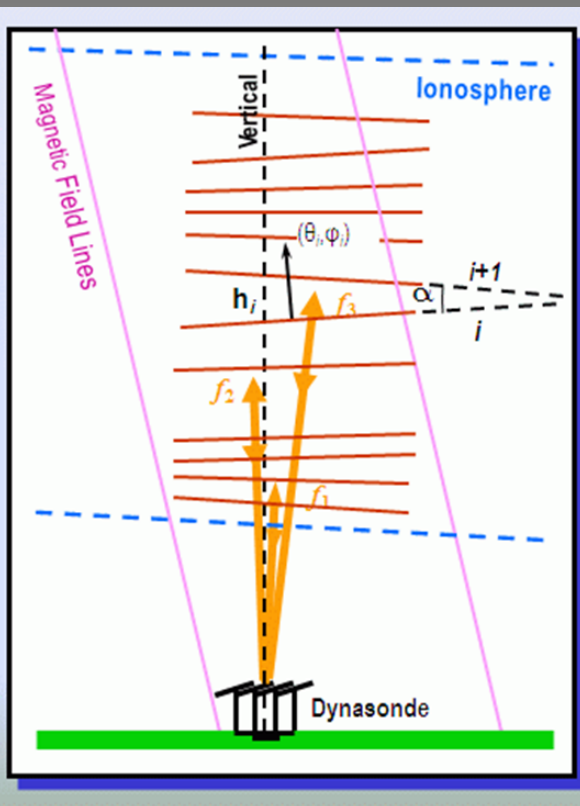
Echo is a physical object characterized by 7 physical parameters (two angles of arrival, group range, Doppler, polarization, phase range, amplitude) plus their uncertainties. Processing the list of the echoes instead of traditional amplitude-based image analysis is the distinctive property of a Dynasonde system. This technique turns radar into a measuring system, not merely imaging system.

Feature important for Acoustic Gravity Wave studies:

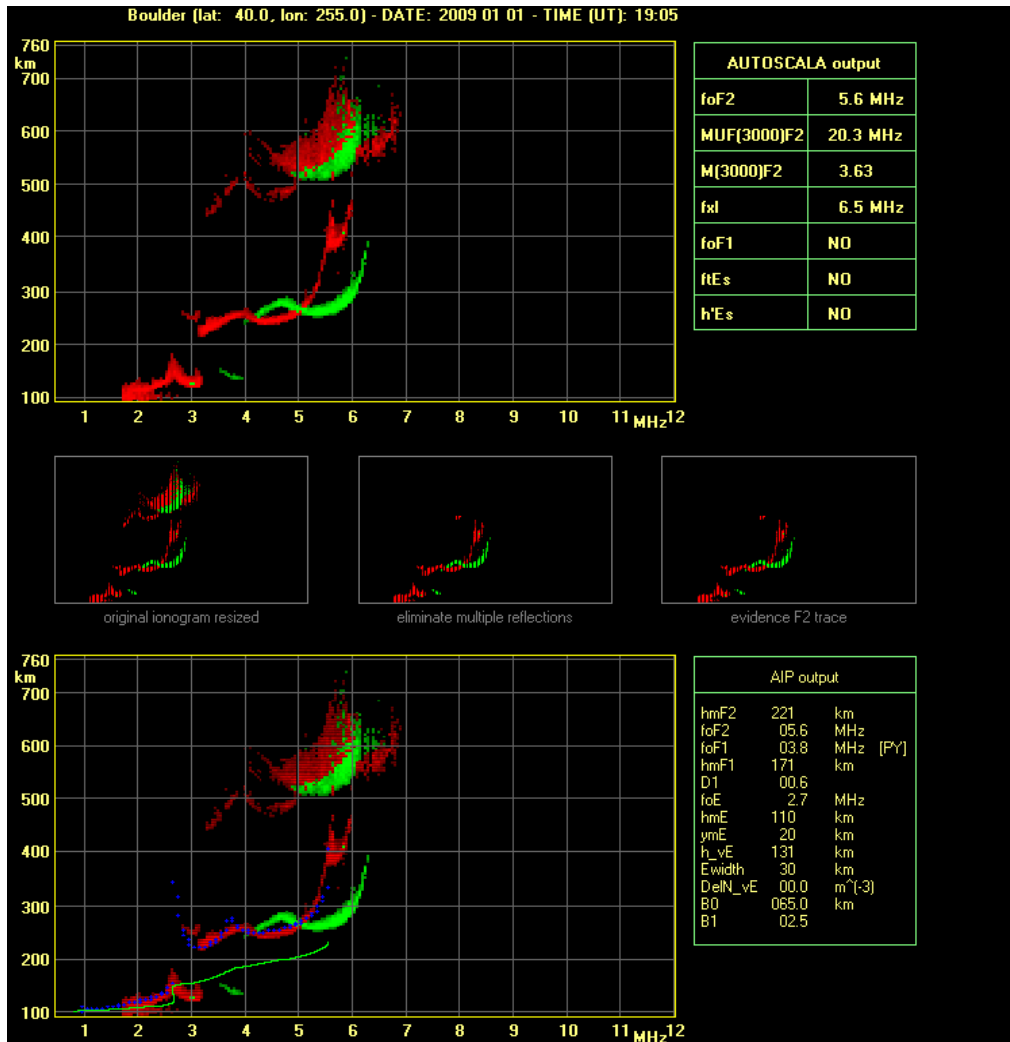
Precision measurements of the range and the angles of arrival for every echo.

NeXtYZ (“Next Wise”), 3-D Plasma Density Inversion Procedure

provides parameters of the Wedge Stratified Ionospheric Model. [Zabotin et al., Radio Sci., 2006]

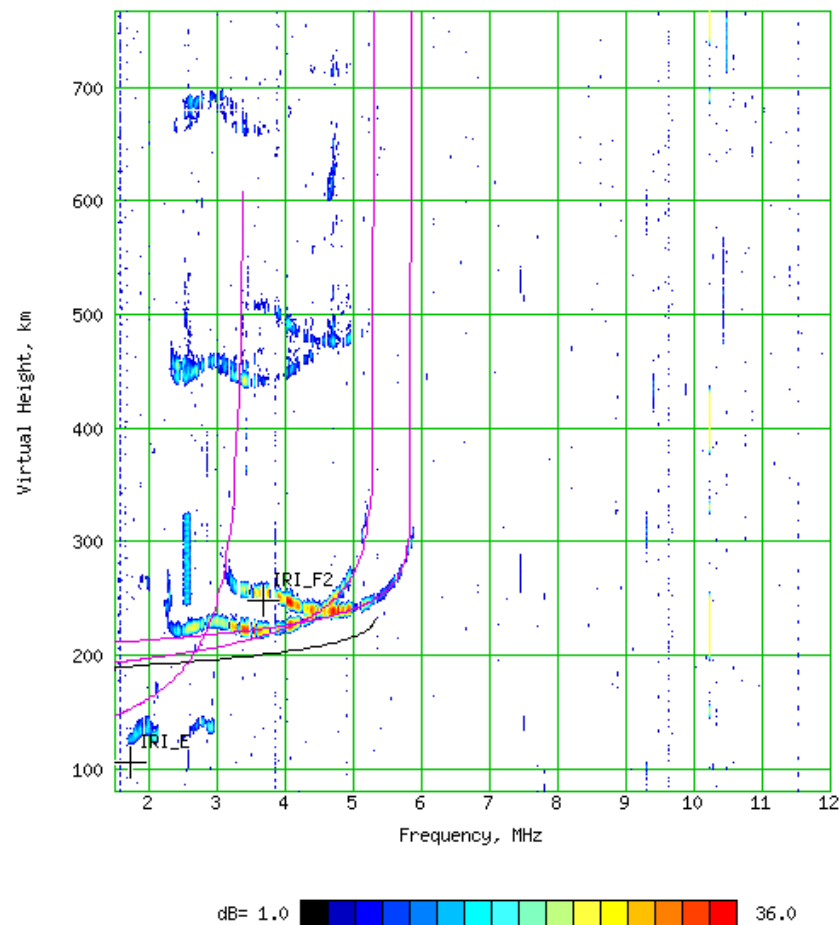


Automatic Scaling Options



Autoscala from INGV

Hanscom Vertical Incidence Pulsed Ionospheric Radar (HAJ45) ESIR Ionogram
Signal-to-Noise at 2010/03/20 (079) 11:42:02 UT (06:42:02 LST)



spacenv.com



ESIR

ESIR from SEC

Contact these vendors for terms and conditions of use for this software

Frequency Allocation

Ionosonde is an active transmitting system

- RF License needed

Ionosonde frequency use is non-standard

- Regulatory agencies used to HF comms allocations

- Exclusive use of narrow band channels

Antenna radiation pattern is Vertical Incidence

- Little RFI beyond a few km

Some 200 ionosondes have operated since 1930

- Precedents exist

A bad frequency allocation can cripple the instrument

Frequency Allocation Issues

Ionosondes do not fit into standard frequency allocations

20 kHz bandwidth vs 5 kHz allocations

Exclusive Use vs Temporary Access

Occupies any 5 kHz channel 1 second every minute

Complete MF-HF Spectrum coverage required

At least 1-16 MHz

A few specific narrowband channels can be excluded

Standard time frequencies (2.5, 5.0, 10.0, 15.0, 25.0 MHz)

Exclusion of whole bands is disastrous

Aeronautical, Marine, Land Mobile, Fixed Terrestrial

**Site-Specific allocations that address
specific RFI issues are required!**

Transmitter Technical Data

Frequency Range: 0.5 to 26 MHz

4 kW peak power, 5% duty cycle max. 200W average.

Logarithmically spaced frequencies (typical 0.5% spacing)

4 to 16 pulses per frequency, 100 Hz rate

70 microsecond (15 kHz bandwidth) pulses

Raised cosine pulse: Low out-of-band emission

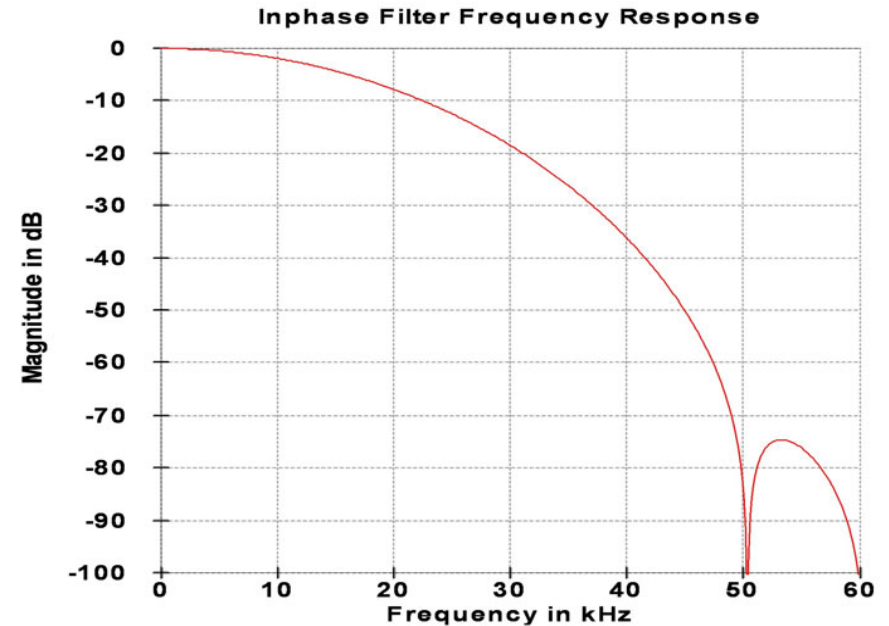
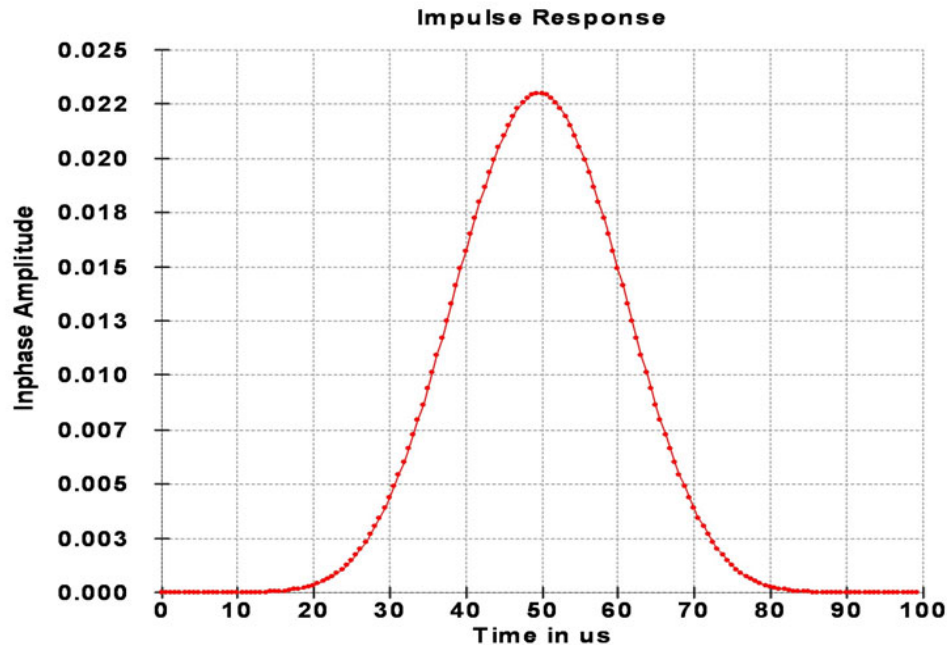
Very low harmonics

Typical 0.1 second channel use per frequency

1 to 5 minute repeat rate

Total channel use: 0.1%: Low probability of RFI

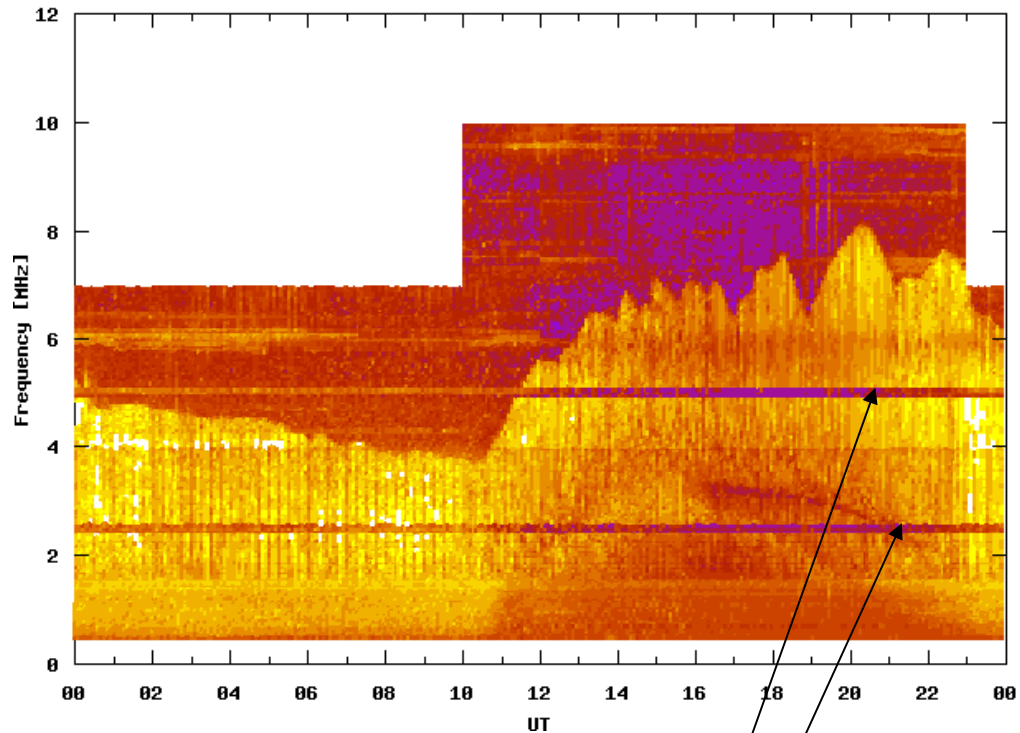
VIPIR Waveform



The receiver impulse response of a raised cosine to the 4th power provides good balance between range resolution and channel occupancy without substantial artifacts in either the time or frequency domain

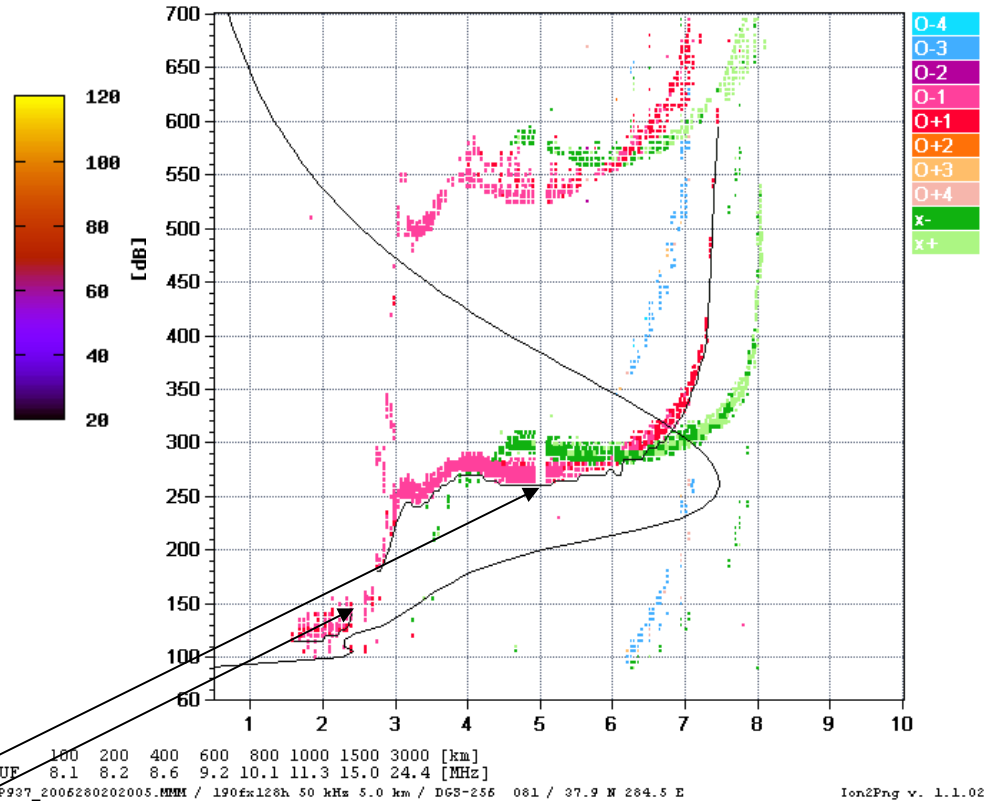
Success: Wallops Island

Max Signal Plot for WP937_2006280



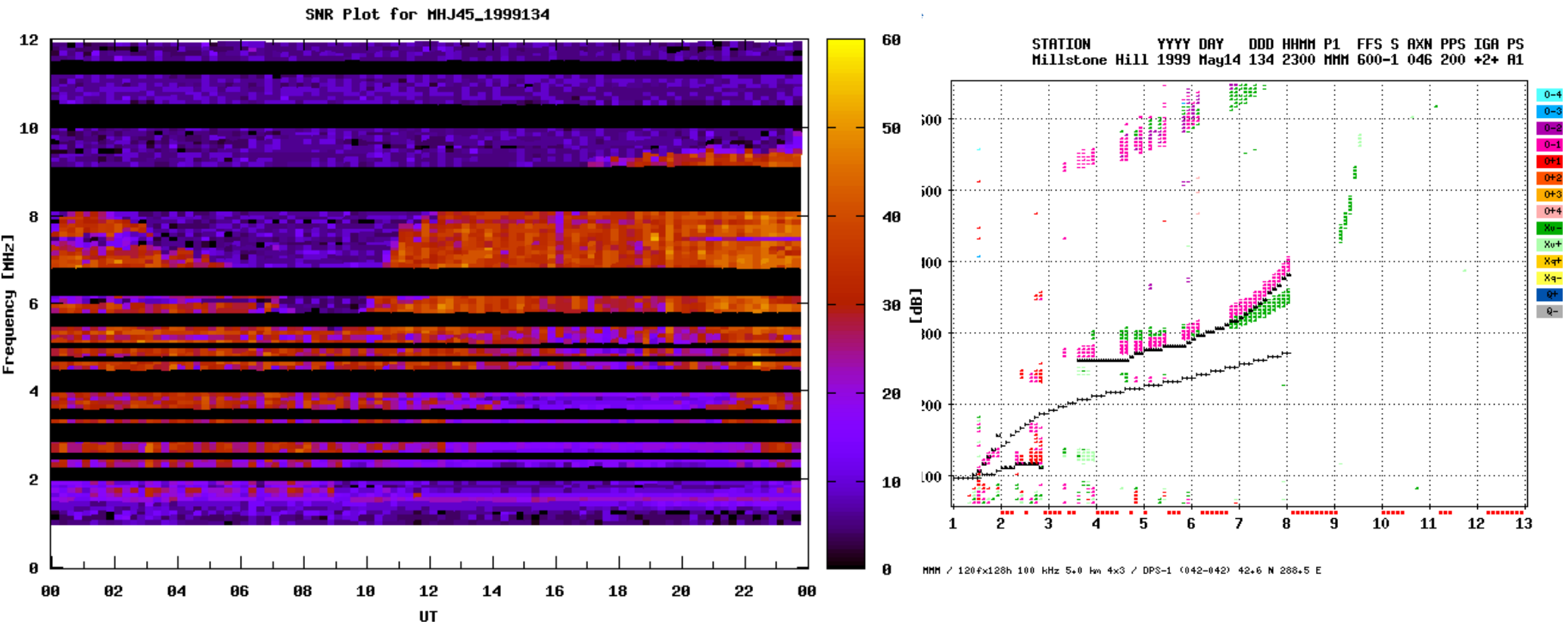
Lowell

Station YYYY DAY DDD HHMM P1 FFS S AXN PPS IGA PS
Wallops Island 2006 Oct07 280 2020 MMM 1 045 200 33+ A1



Minor restrictions mitigate RFI while maintaining data integrity

Failure: Millstone Hill



Major restrictions destroy the integrity of the data

Results: Boulder Digisonde vs VIPIR



Station YYYY DAY DDD HHMM P1 FFS S AXN PPS IGA PS
Boulder 2008 Dec31 366 1630 MMM 1 046 200 32+ A1

Improved
Measurement Clarity

foF2 4.575
foF1 3.41
foF1p N/A
foE 2.31
foEp 2.41
fxI 5.45
foEs 3.00
fmin 1.80

MUF(D) 15.60
M(D) 3.43
D 3000.0

h'F 205.0
h'F2 221.0
h'E 115.0
h'Es 123.0

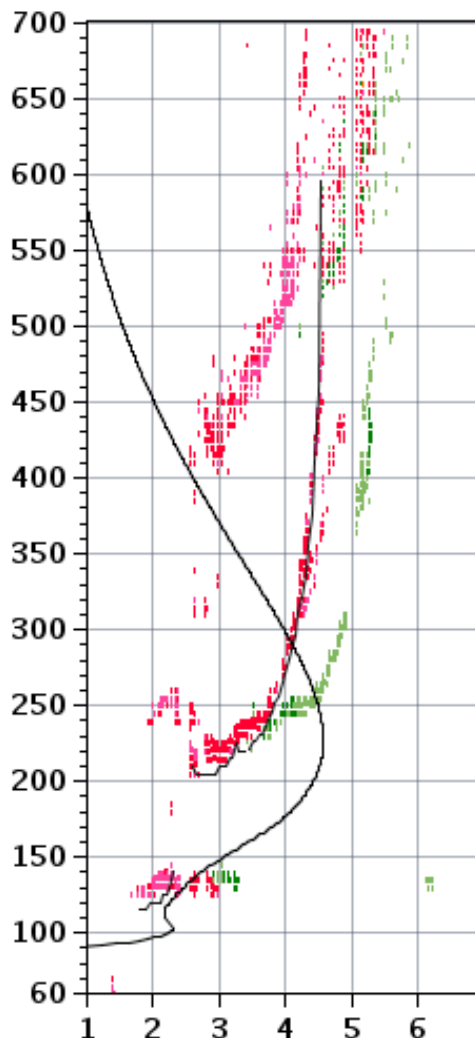
hmF2 225.5
hmF1 159.5
hmE 103.3

yF2 99.0
yF1 22.6
yE 13.1

B0 102.5
B1 2.40

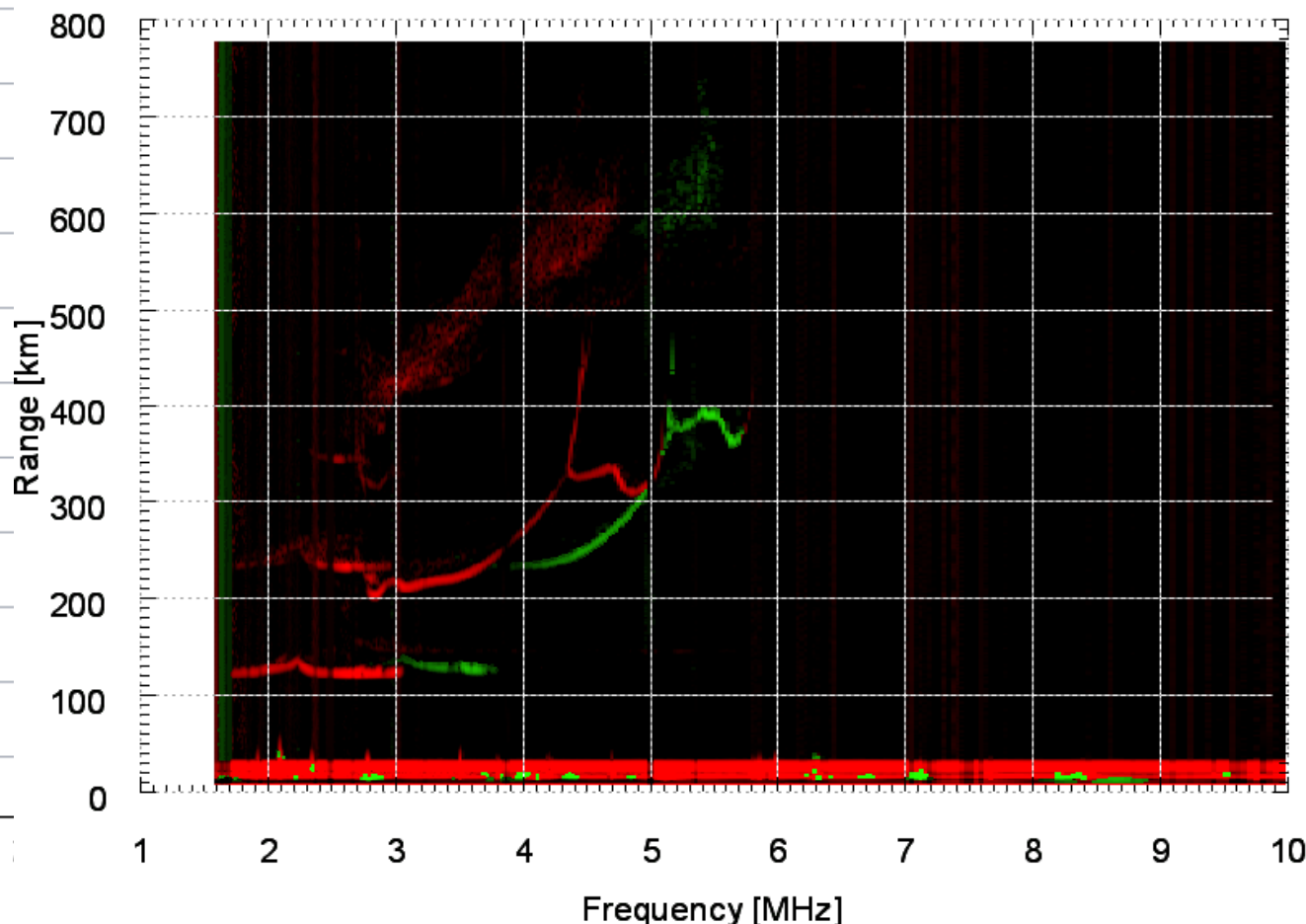
C-level 21

Auto:
Artist4.5
200311



BD840_2008366163519

O&X SNR [dB]



D 100 200 400 600 800 1000 1500 3000 [km]
MUF 5.2 5.3 5.5 5.9 6.5 7.3 9.6 15.6 [MHz]
BC840_2008366163005.MMM / 300fx128h 50 kHz 5.0 km / DGS-256



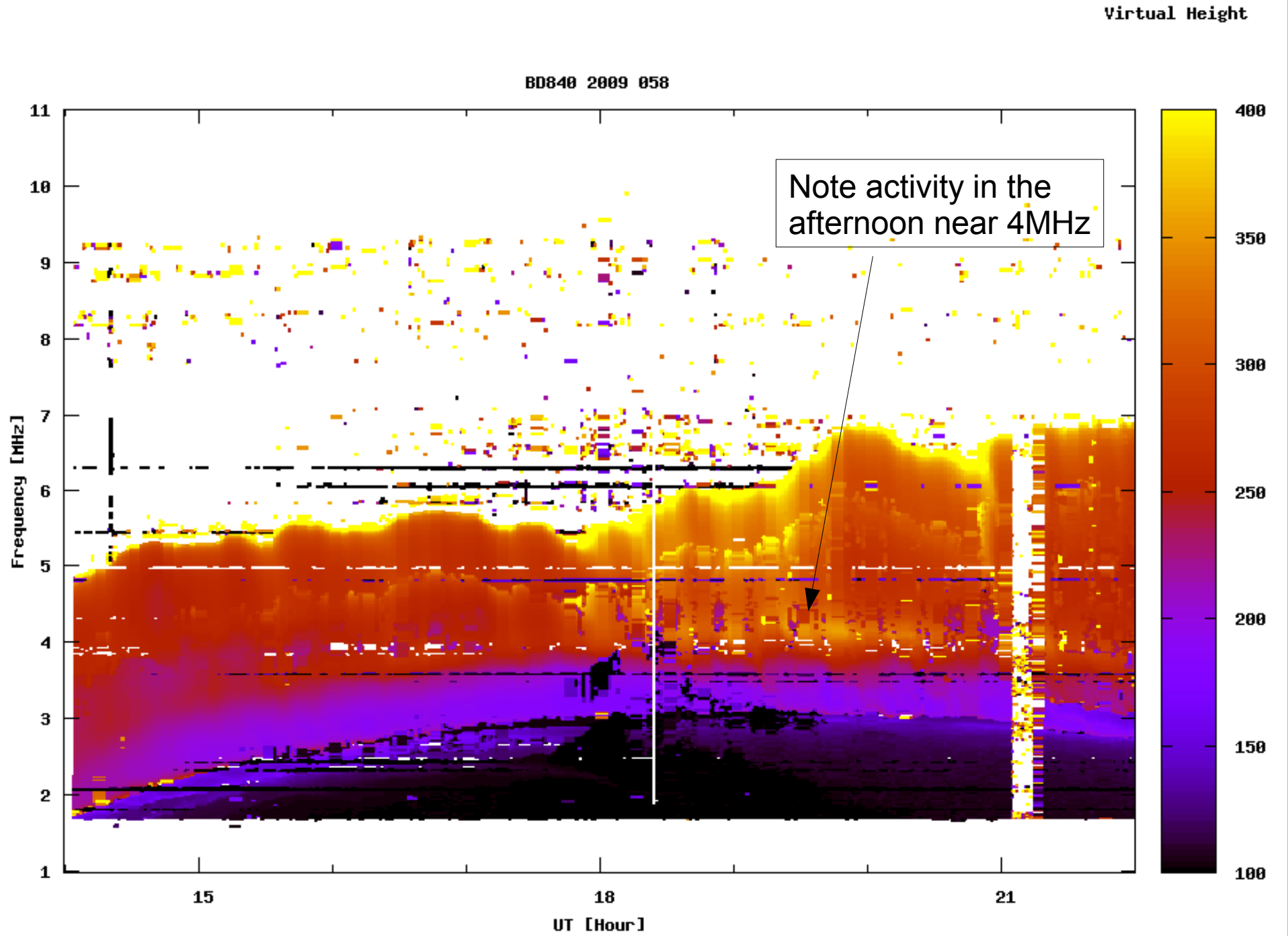
0 5 10 15 20 25 30 35 40

0 5 10 15 20 25 30 35 40

O

X

Boulder 1-minute Data

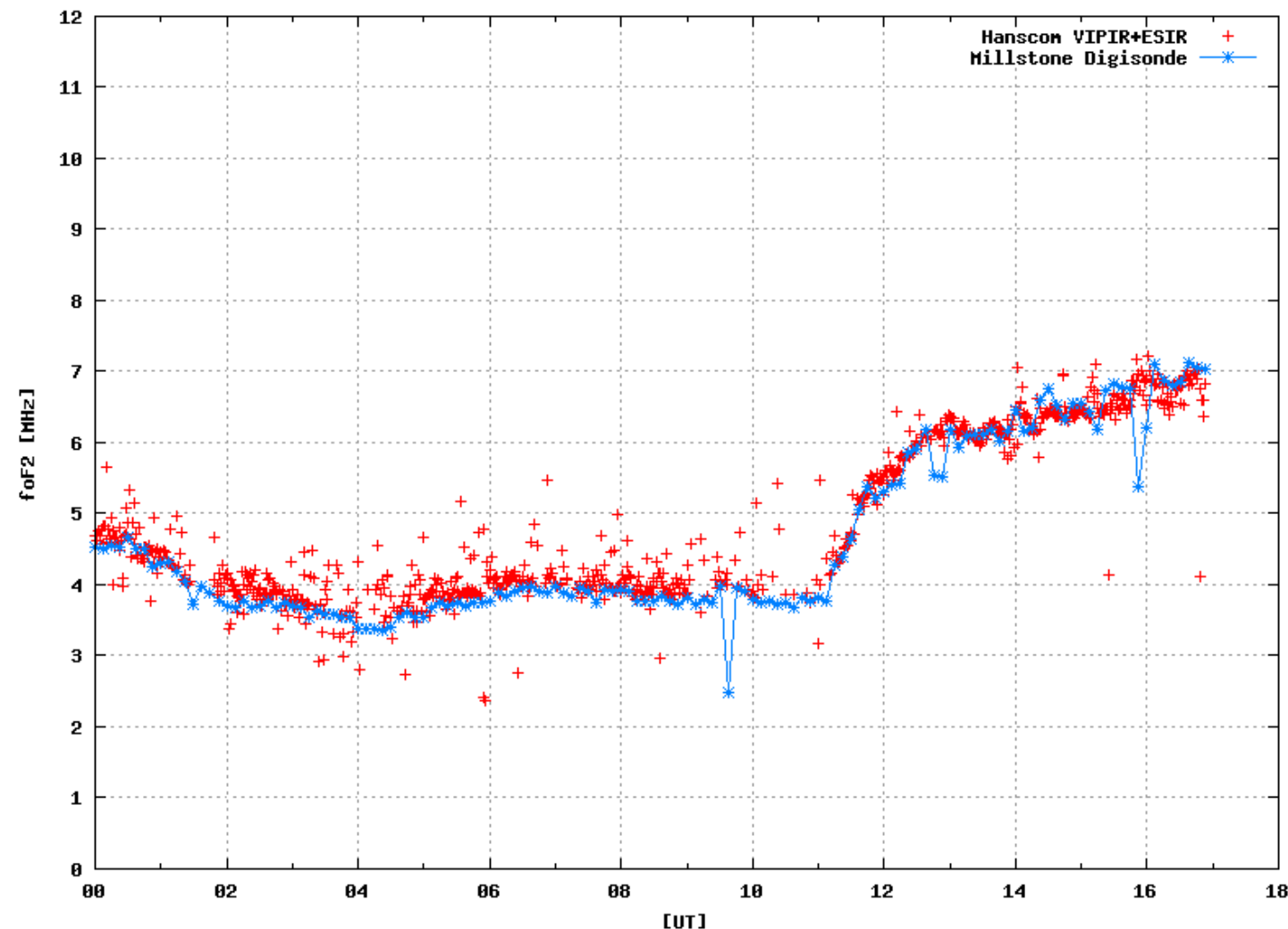


High Time Resolution Data

VIPIR can comfortably make 1 ionogram per min

Digisondes are experimenting similarly

foF2 comparison HANSCOM and MILLSTONE 22Feb18



Plot of foF2 for 17 hours from the Millstone Hill Digisonde (Blue) and Hanscom VIPIR (Red)

Data courtesy of U.Mass Lowell and Boston College

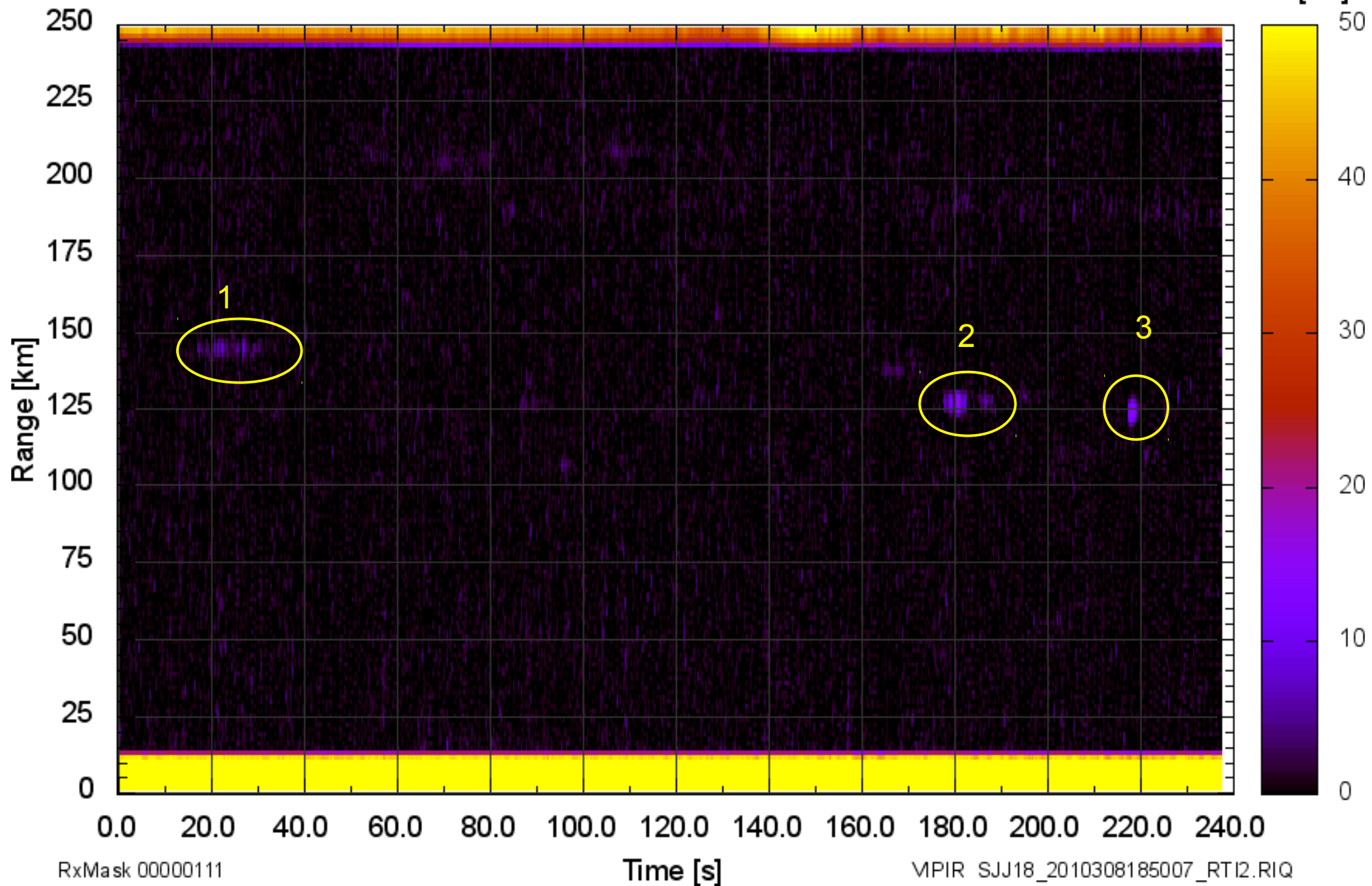
How to optimally use these data?

Meteor Trails at 5.8 MHz

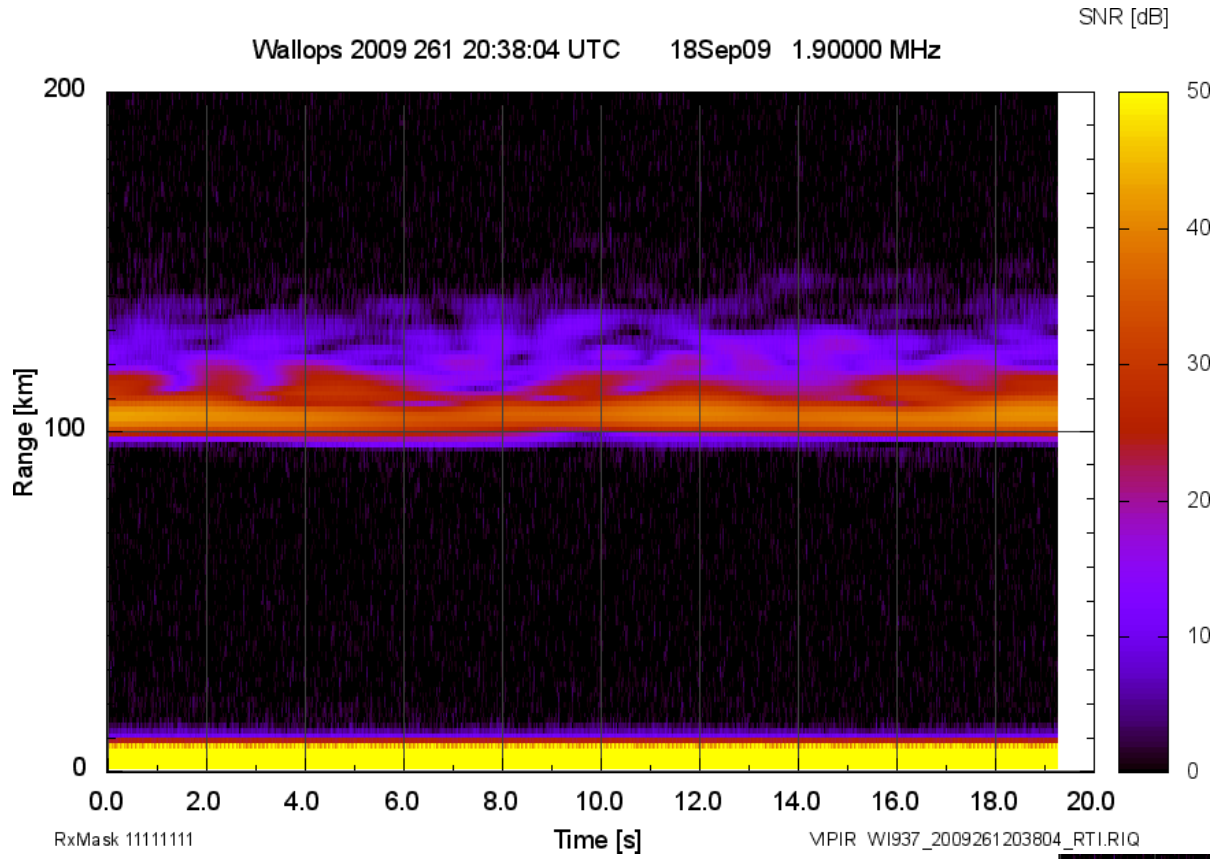
San Juan 2010 308 18:50:07 UTC

04Nov10 0.00000 MHz

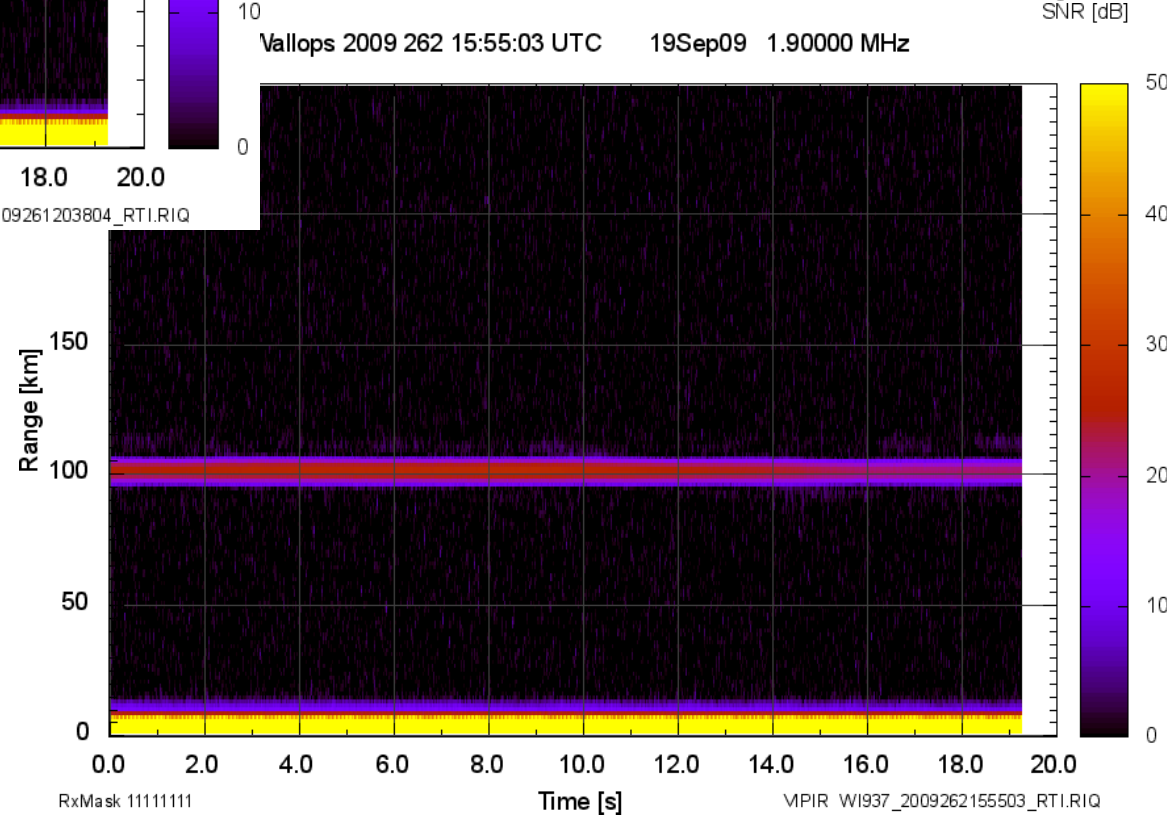
SNR [dB]



Plasma Turbulence



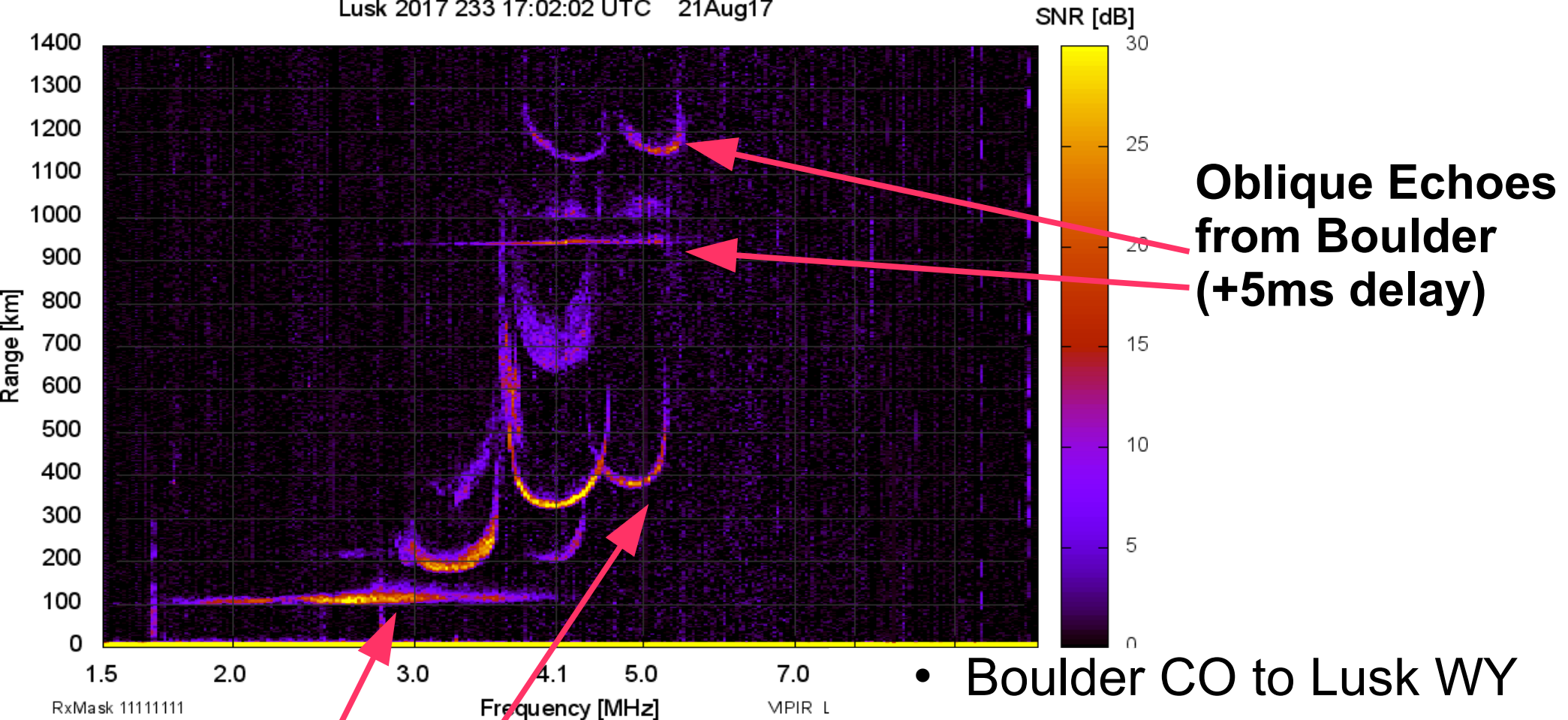
Smooth E- layer



Structured E- layer

Oblique Propagation – Solar Eclipse 2017

Lusk 2017 233 17:02:02 UTC 21Aug17

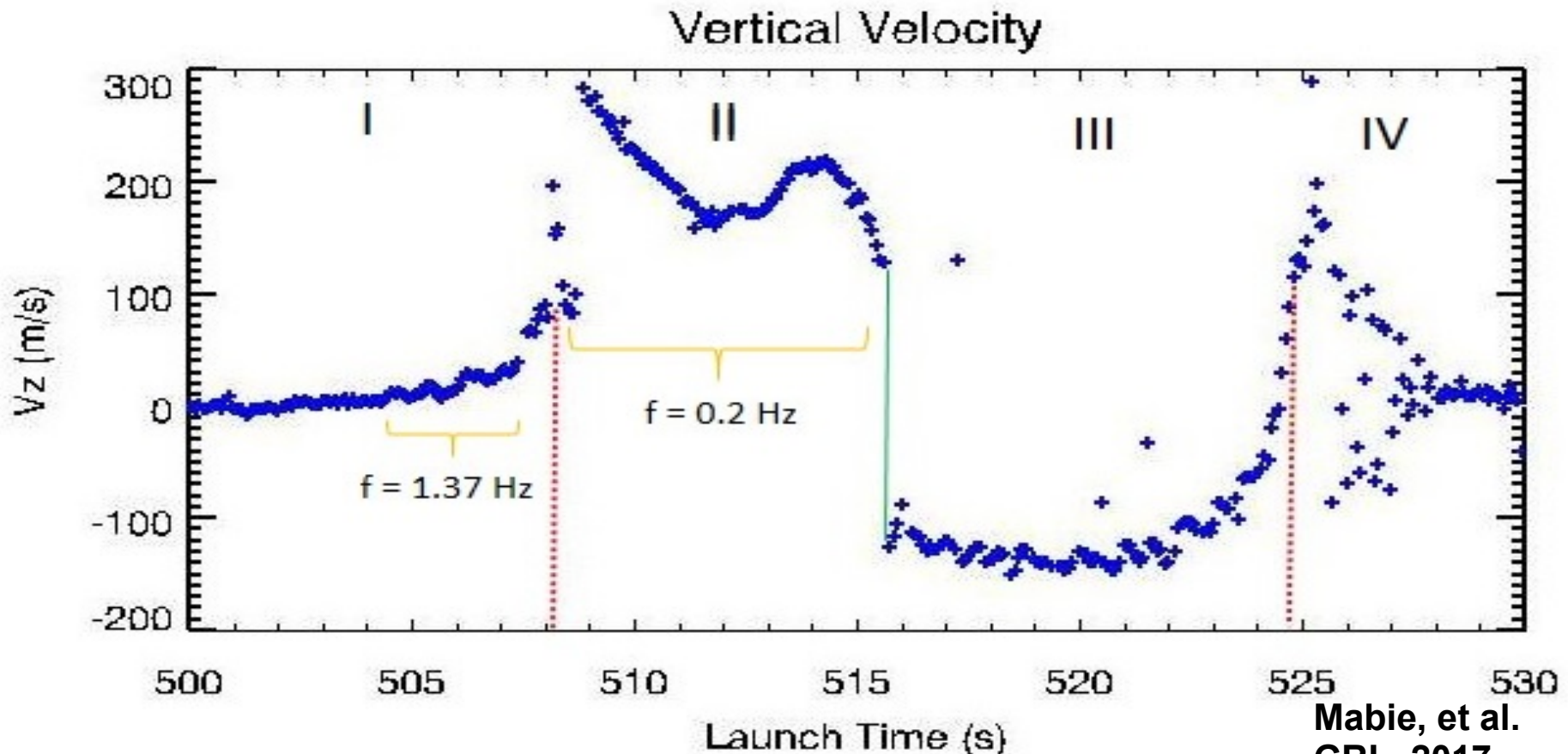


**Vertical Echoes
from Lusk**

**Oblique Echoes
from Boulder
(+5ms delay)**

- Boulder CO to Lusk WY
- 280 km ground distance
- GPS disciplined rubidium oscillators
- Simultaneous oblique and vertical sounding

Acoustic waves



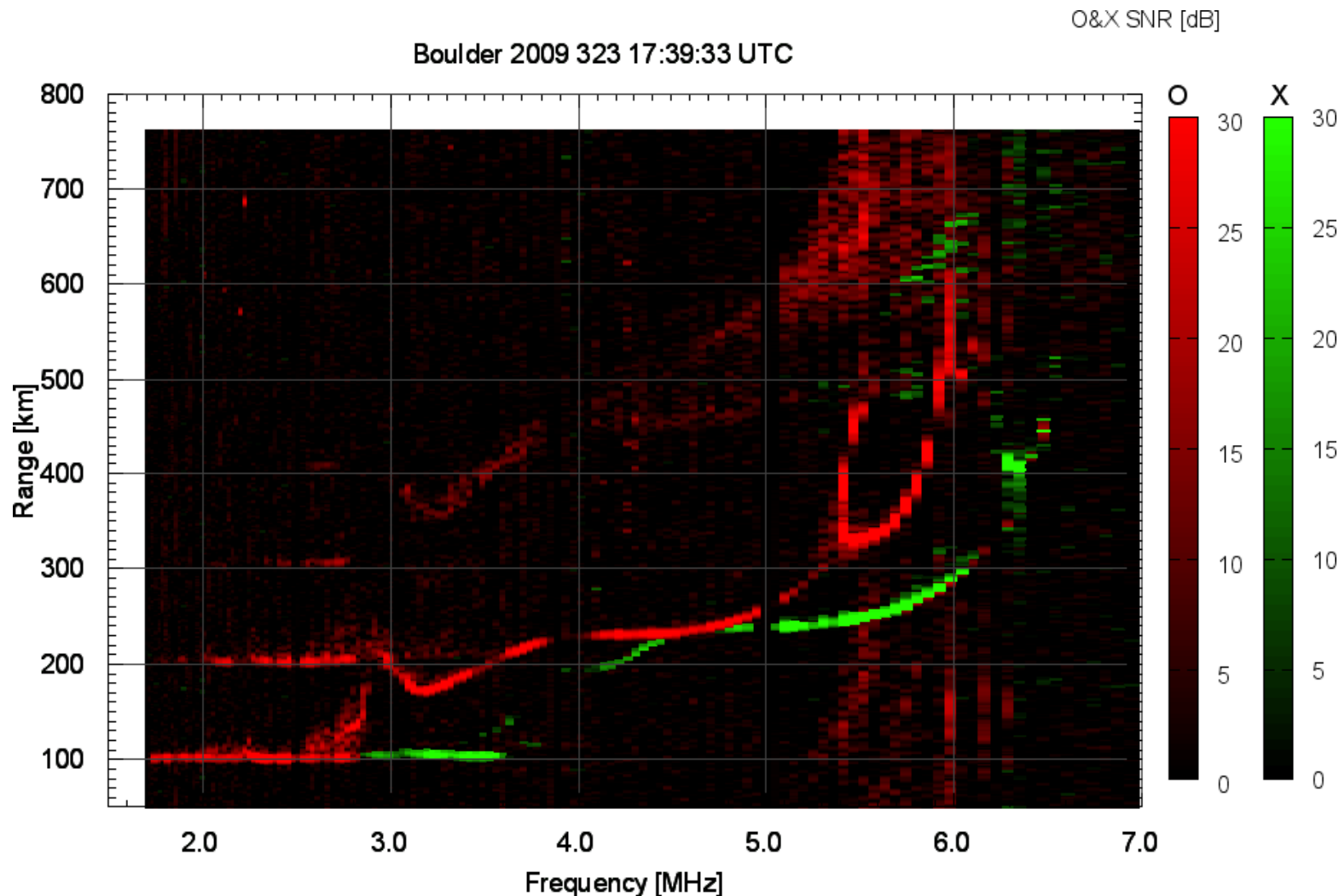
Mabie, et al.
GRL, 2017

- Sound waves from rocket launches perturb the ionosphere
- Possibly similar effects from earthquakes and tsunamis
- Potential to measure sound speed in the thermosphere

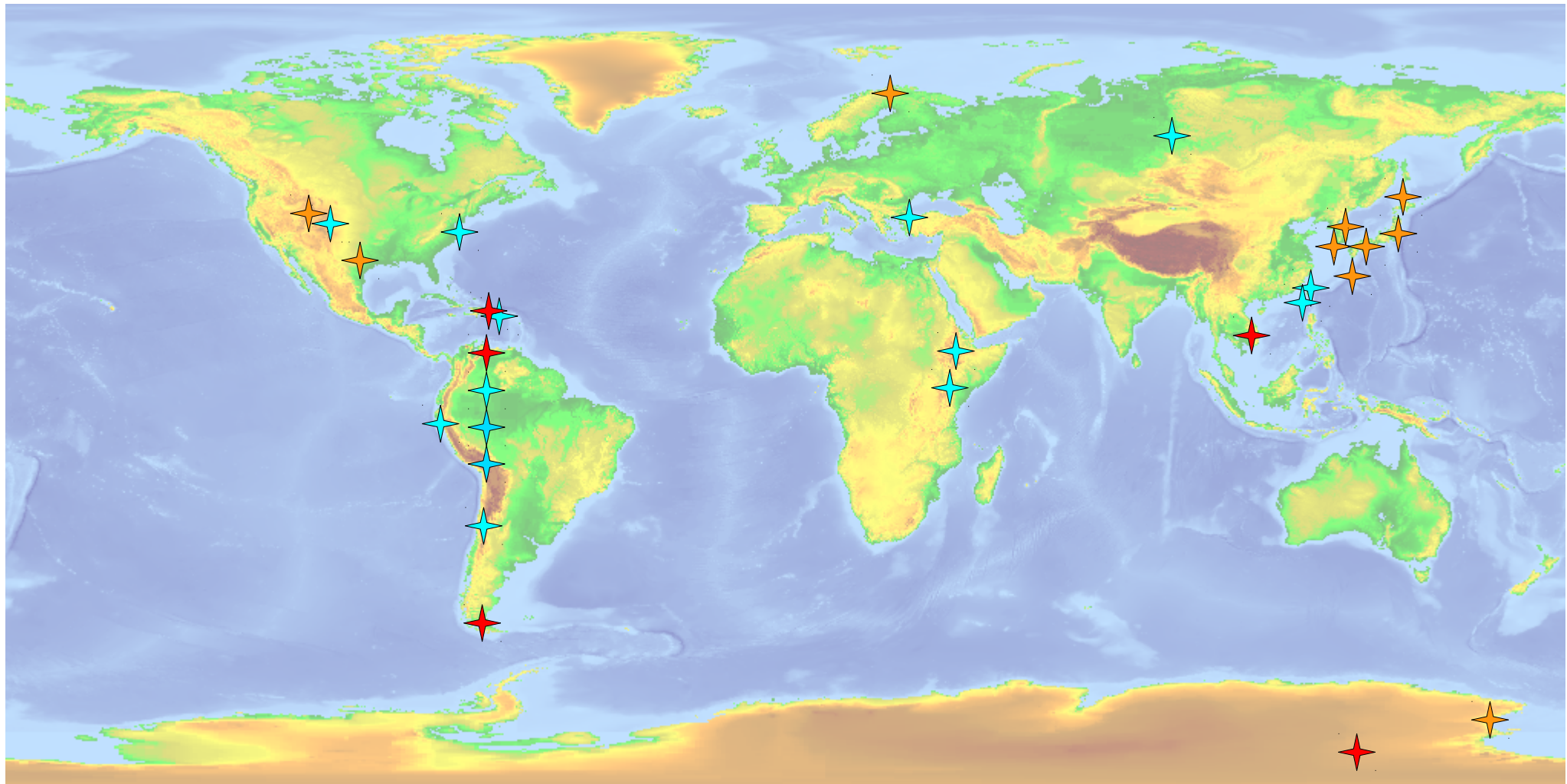
Very Fast Sweeps

Ionogram sweeps < 10 seconds long

Continuous repeat of 100's of sweeps possible



VIPIR Facilities



★ VIPIR 1 (14)

★ VIPIR 2 (10)

★ Planned (5)

Updated November 2017

Internet Resources

- World Data Center A, Boulder:
<http://www.ngdc.noaa.gov/stp/IONO/ionohome.html>
- Digisondes and ARTIST : <http://ulcar.uml.edu/> <http://www.digisonde.com/>
- Autoscala: <http://roma2.rm.ingv.it/en/facilities/software/18/autoscala>
- ESIR : <http://www.spacenv.com/>
- Low-latitude Ionospheric Sensing System: <http://jro.igp.gob.pe/lisn/>
- Vertical Incidence Pulsed Ionosphere Radar (VIPiR): Terry.Bullett@noaa.gov
- Canadian Advanced Digital Ionosonde (CADI): <http://cadiweb.physics.uwo.ca/>
- Ionospheric Prediction Services (IPS): <http://www.ips.gov.au/>
- Ionosonde Network Advisory Group (INAG)
<http://www.ips.gov.au/IPSHosted/INAG/>
- SPIDR: <http://spidr.ngdc.noaa.gov/spidr/index.jsp>
- Gravity Waves: <http://surf.colorado.edu>